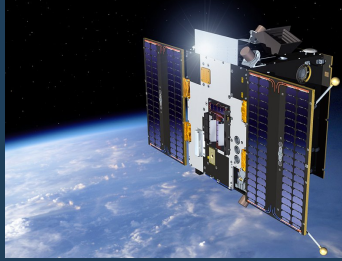


LYRA on-board PROBA2

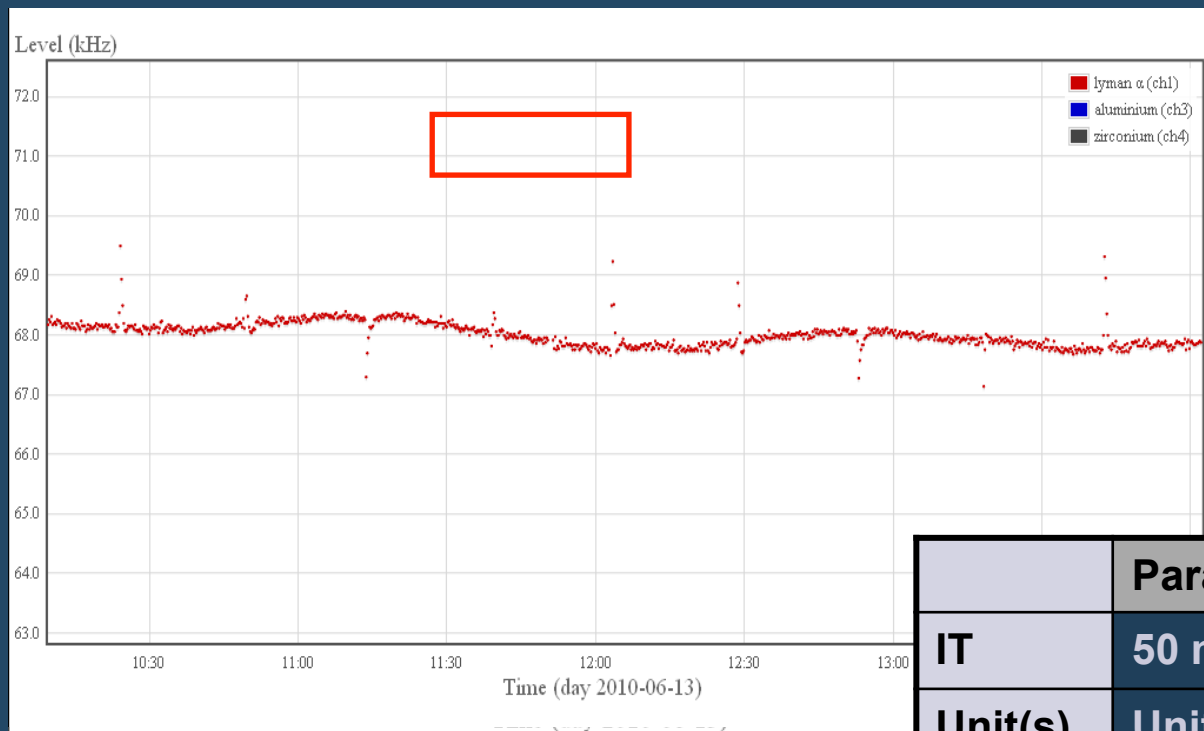
in-orbit operations and
achievements

M. Dominique, J.-F. Hochedez et al.
PROBA2 Workshop
ESTEC, 2010/06/22



Routine activities

<input type="checkbox"/> Nominal acquisition: <ul style="list-style-type: none"><input type="checkbox"/> Unit2<input type="checkbox"/> Integration time = 50 ms	Nominal
<input type="checkbox"/> Calibration <ul style="list-style-type: none"><input type="checkbox"/> Dark current<input type="checkbox"/> LED signal	Weekly
<input type="checkbox"/> Back-up acquisition <ul style="list-style-type: none"><input type="checkbox"/> Acquisition with units1 and 3	Monthly
<input type="checkbox"/> Flat-field analysis <ul style="list-style-type: none"><input type="checkbox"/> Off-pointing sequence	Monthly
<input type="checkbox"/> Bake out - decontamination <ul style="list-style-type: none"><input type="checkbox"/> Switch on of heaters (temperature reaches 50°C)	Once every 6 months



	Parameters of the mode
IT	50 ms
Unit(s)	Unit 2
LED	No
Remark	IT up to 10ms have been tested

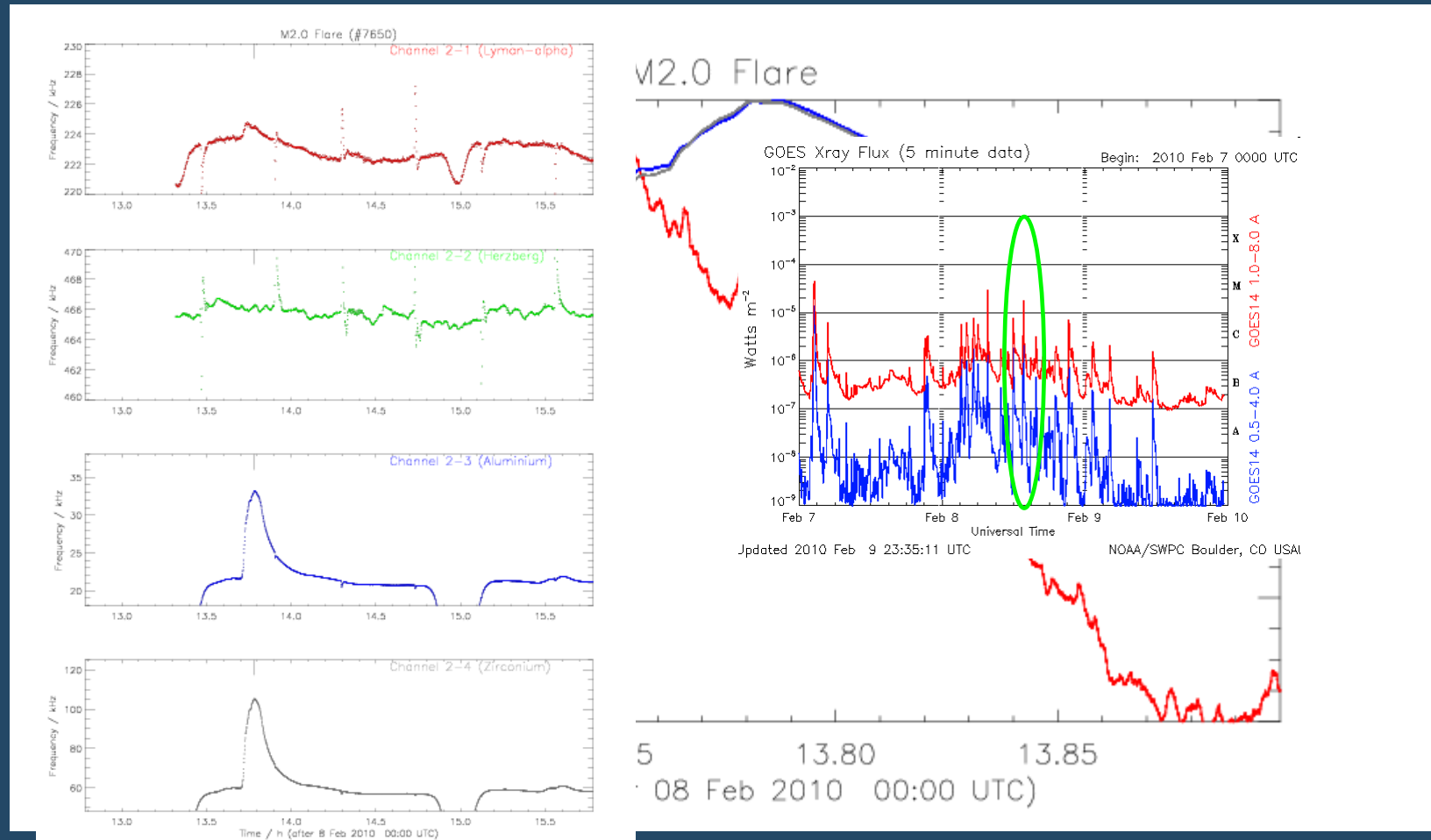


- ❑ Lyra senses flares down to B1.5 at least
- ❑ LYRA list of flares is in agreement with the one of GOES
- ❑ Always visible in the two XUV-EUV channels
- ❑ Some strong and impulsive flares are also visible in Lyman-alpha, which can then be used as a precursor



Flares

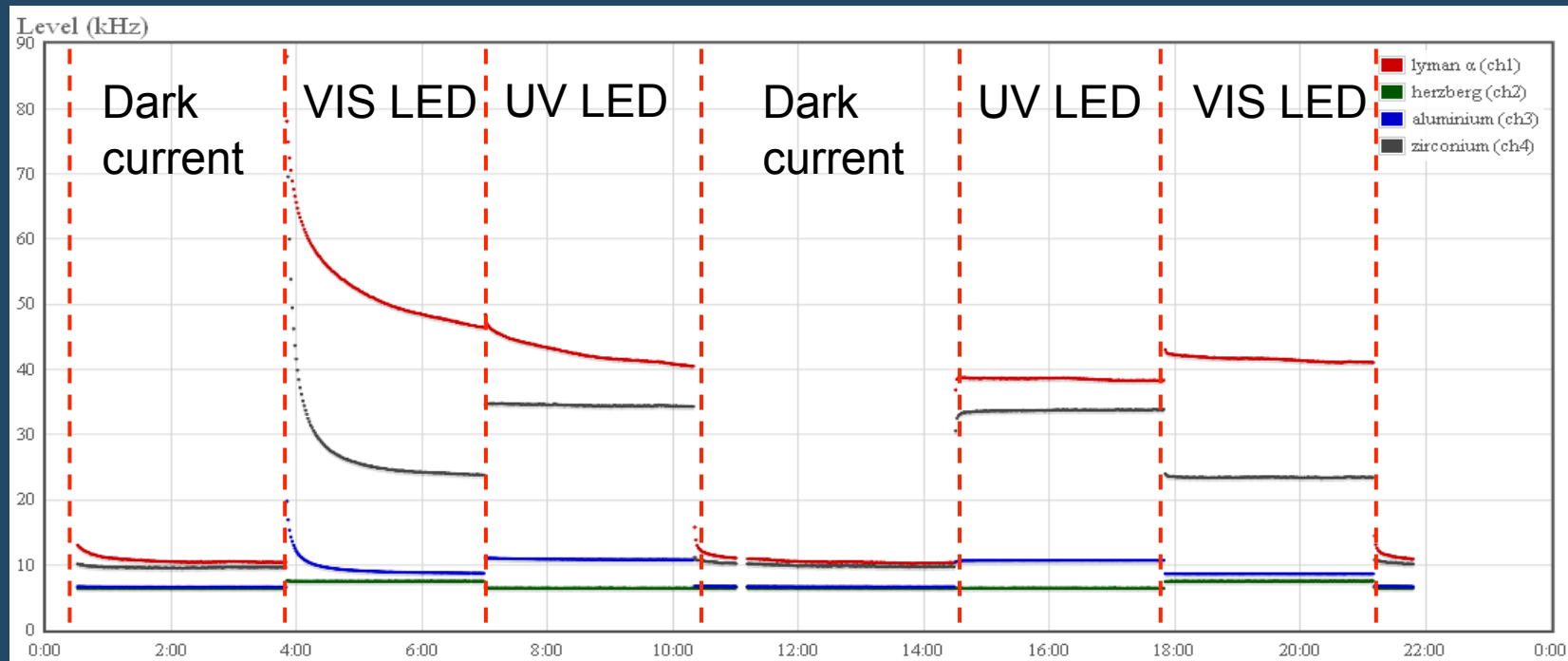
□ M2.0 flare 2010/02/08 – 22h33





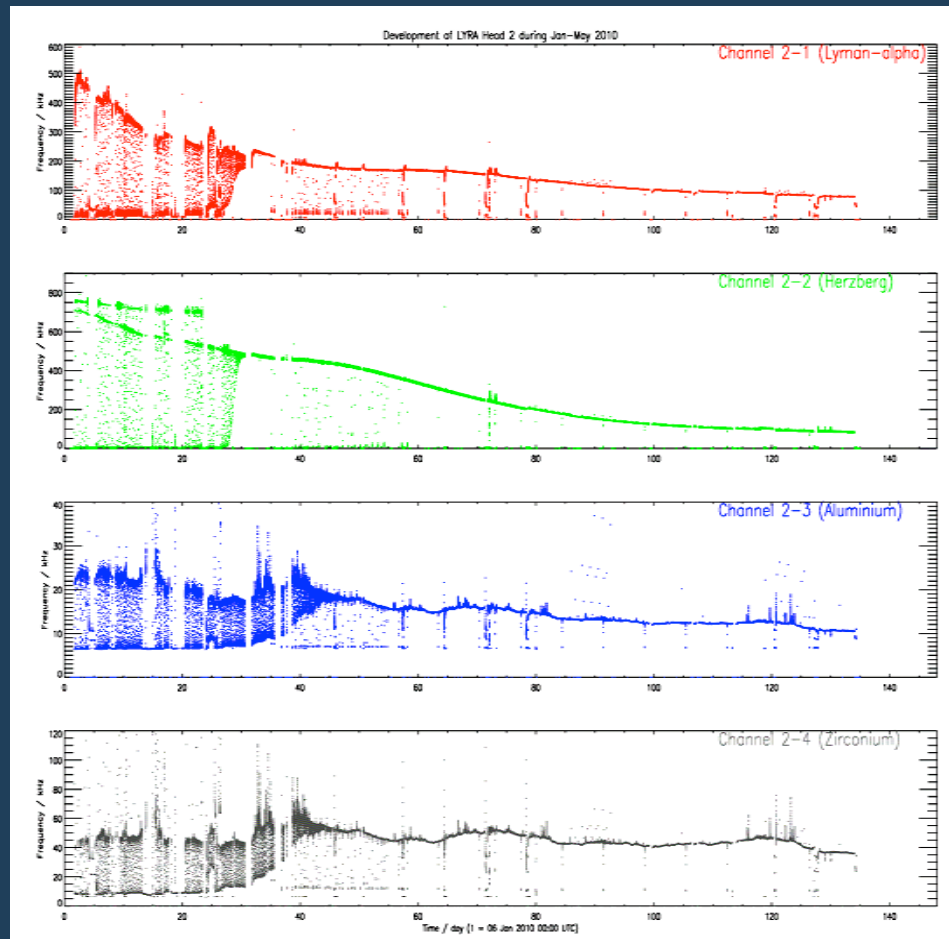
Calibration sequence

- ❑ More than one orbit needed to MSM detectors to reach stabilization.
- ❑ No significant difference between the LED signal at the beginning of the mission and the current one
→ the degradation is affecting the filters and not the detectors

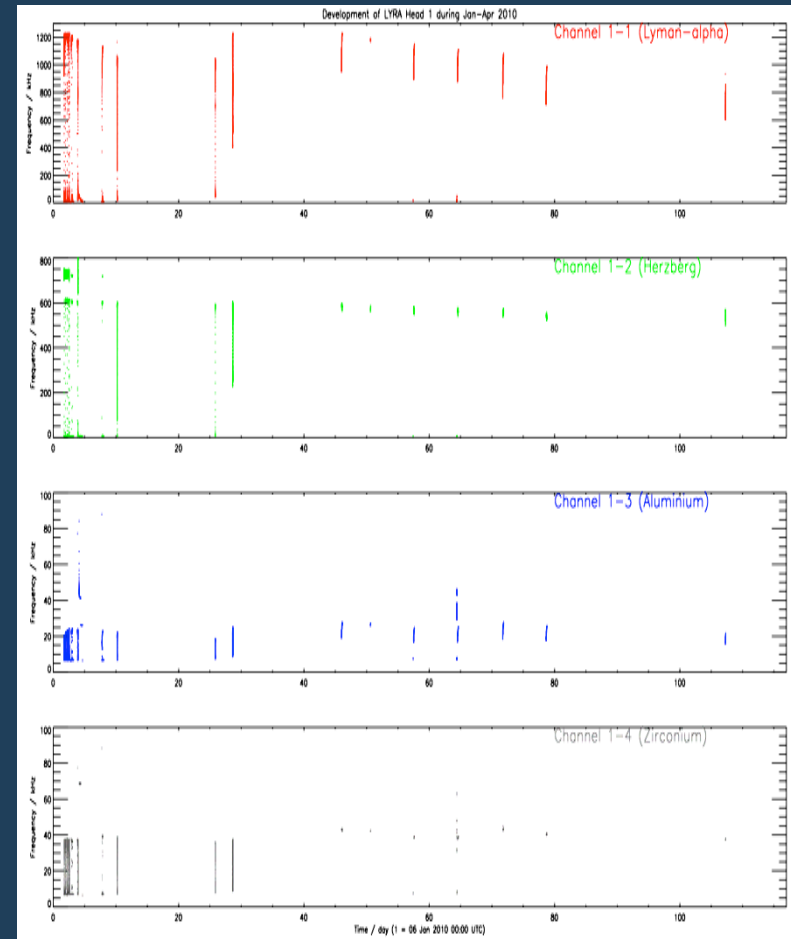




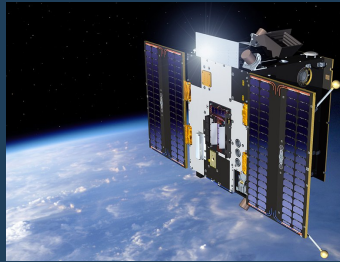
Back-up acquisition



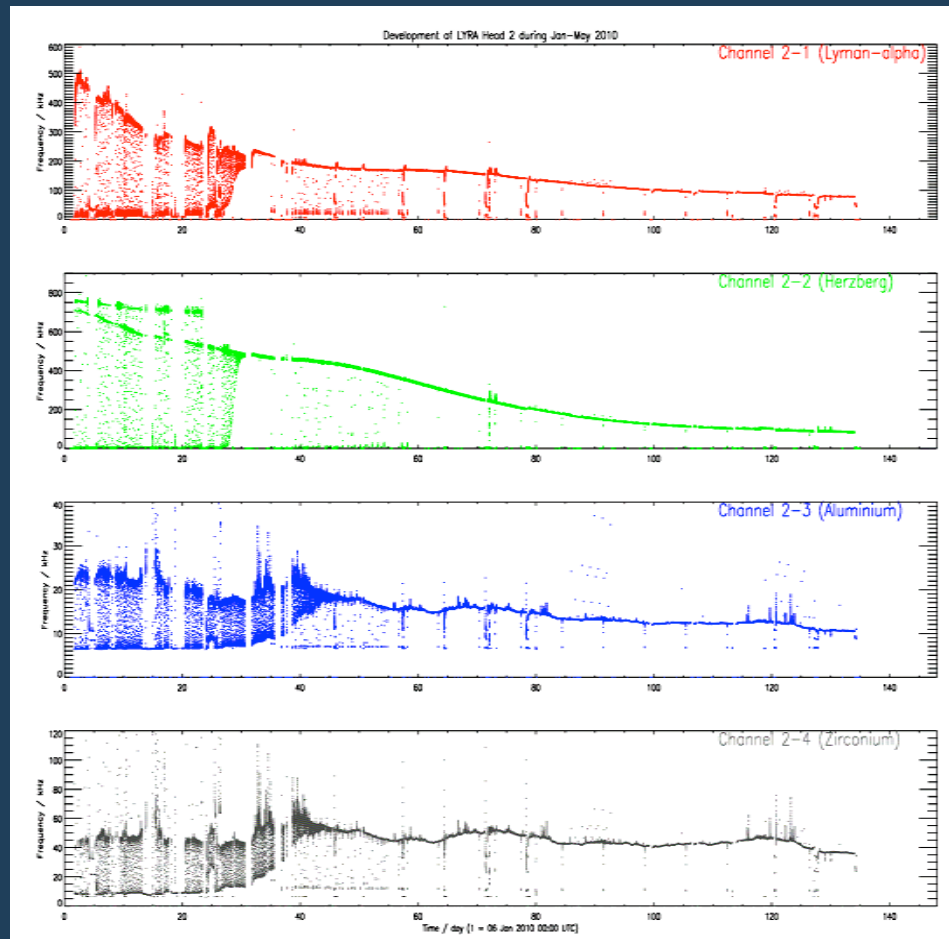
UNIT 2



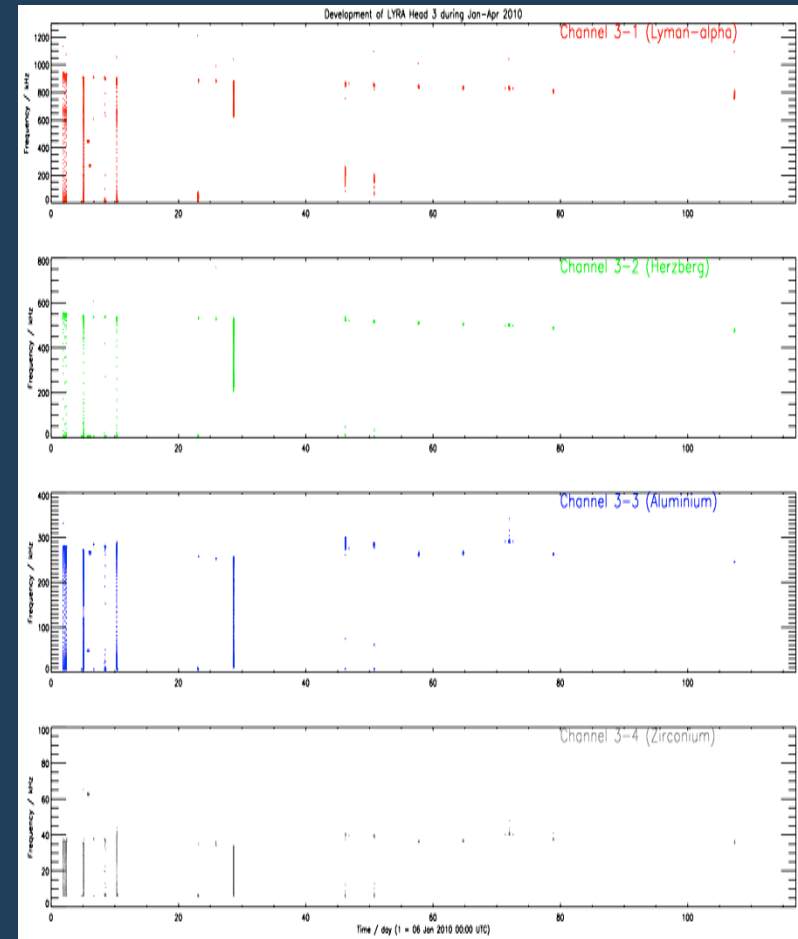
UNIT 1



Back-up acquisition



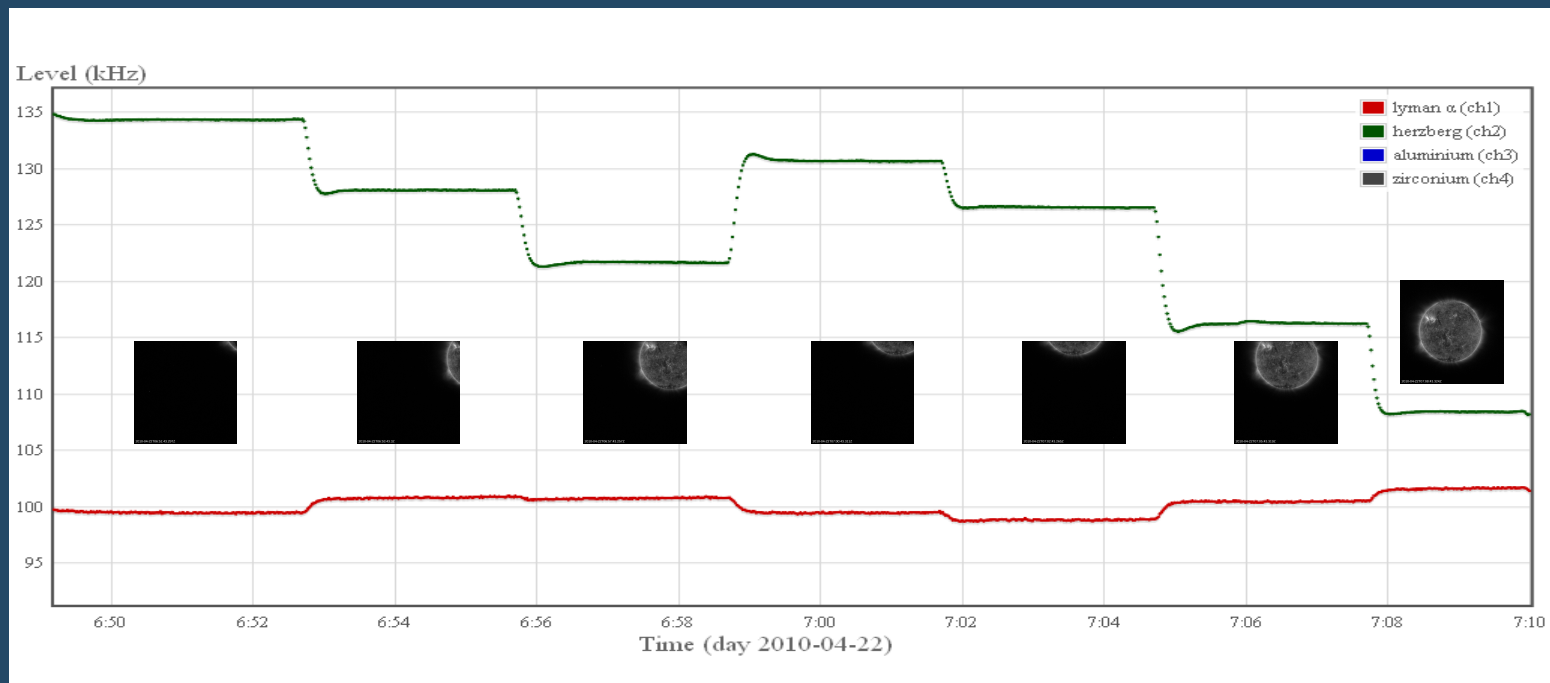
UNIT 2



UNIT 3

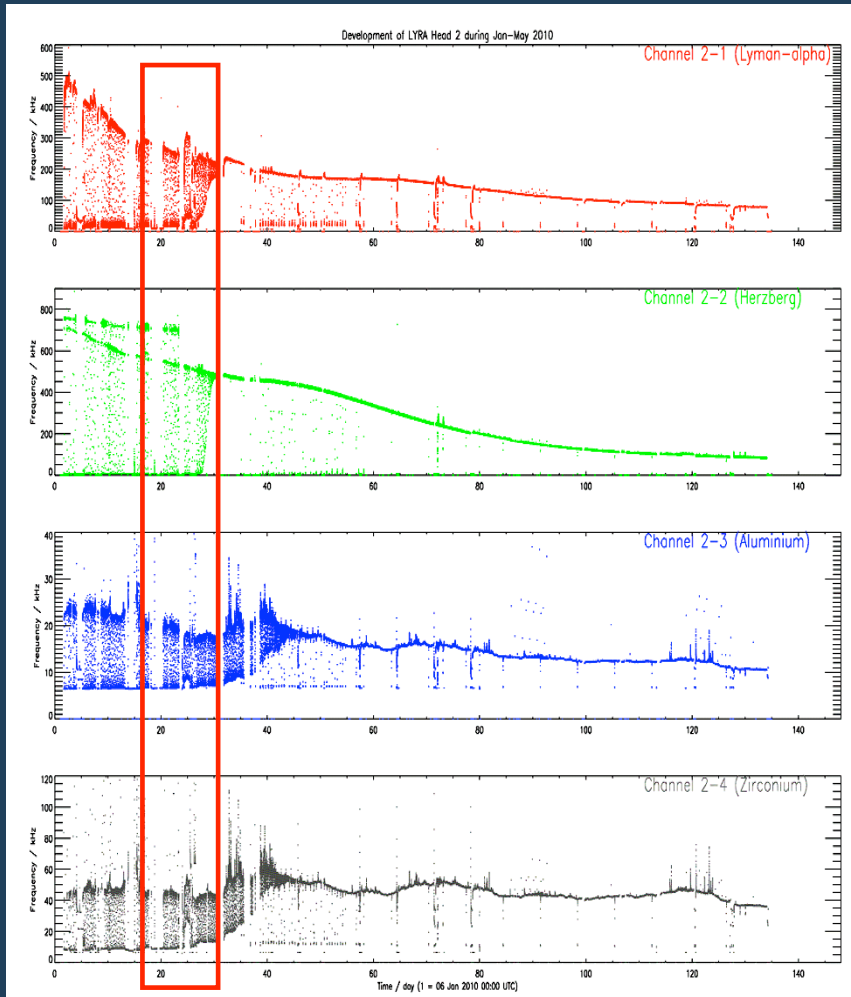


- Paving step : 0.3°
- The Hz channel signal doesn't fit the modeled behavior based on pre-launch flat-field measurements
⇒ slight imprinted degradation ???





Bake-out



- AB + CD heaters on for 24h
- Temperature increase of 10°

No significant effect on the data
Next attempt: July 01



Specific campaigns

- ❑ Occultations
- ❑ Sun-Moon eclipses
- ❑ Stray light analysis
- ❑ SDO/EVE cross-calibration campaign

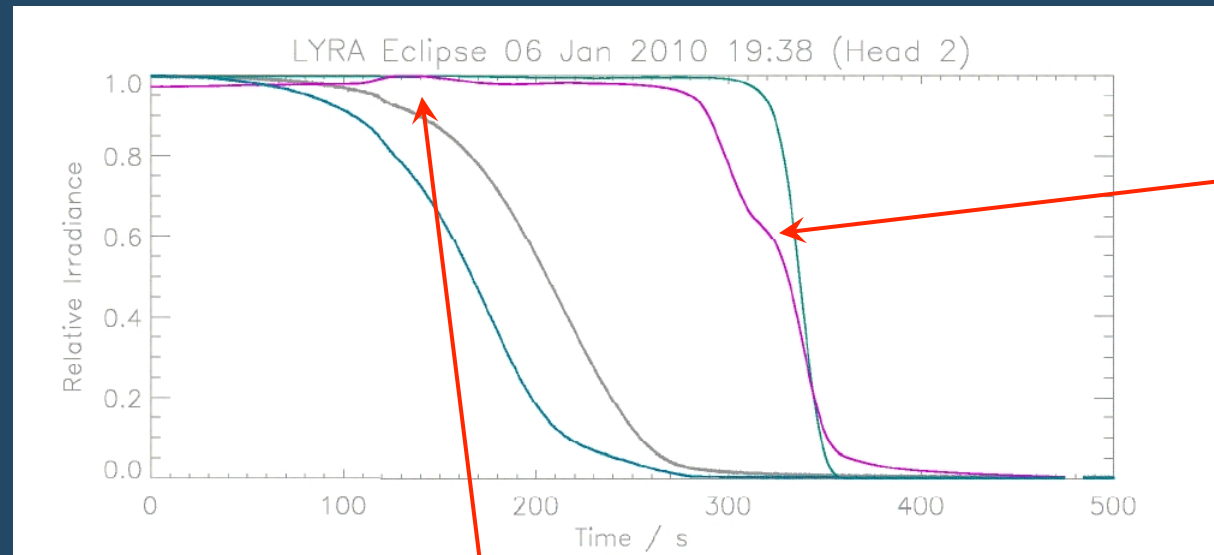
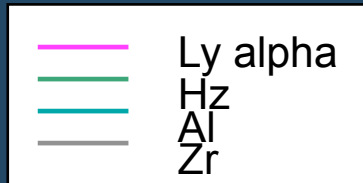
Every orbit till March
January 15 (annular from LYRA)
July 11 (partial from LYRA)

March 18-19
(still under analysis)

May 3
(still under analysis)



Occultations

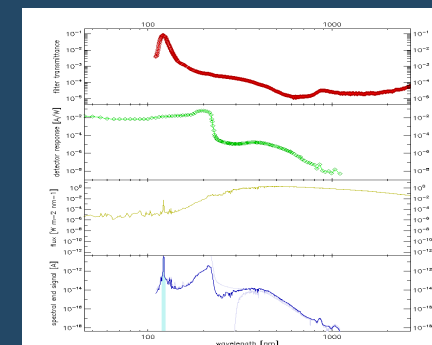


Attitude change

Lyman alpha
completely
absorbed

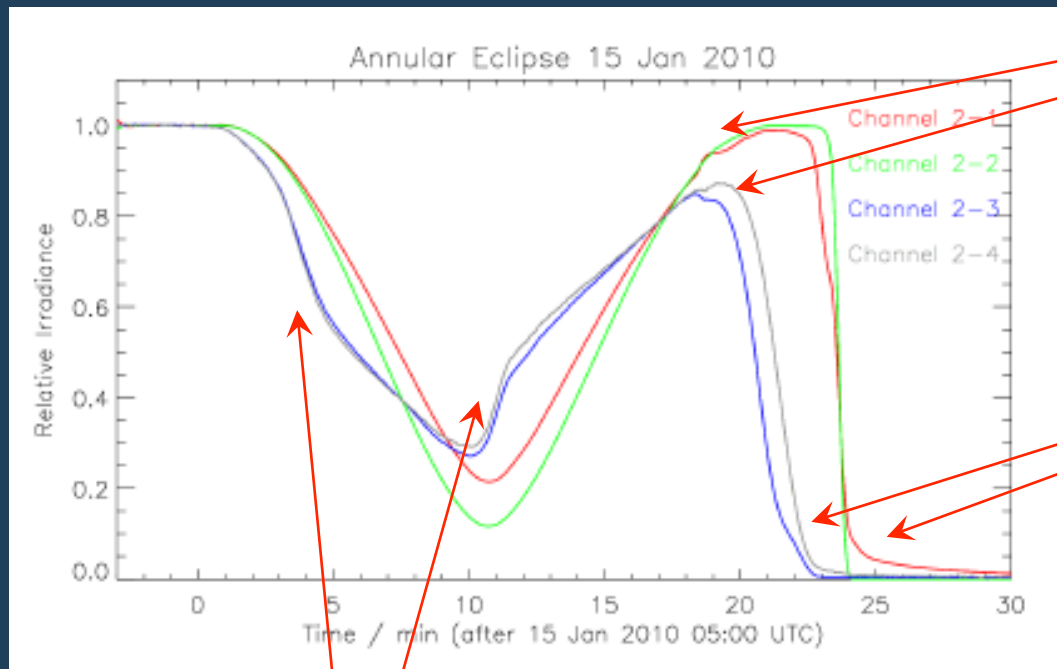


Contribution of
longer
wavelengths





Sun-Moon eclipse



Attitude change

Occultation

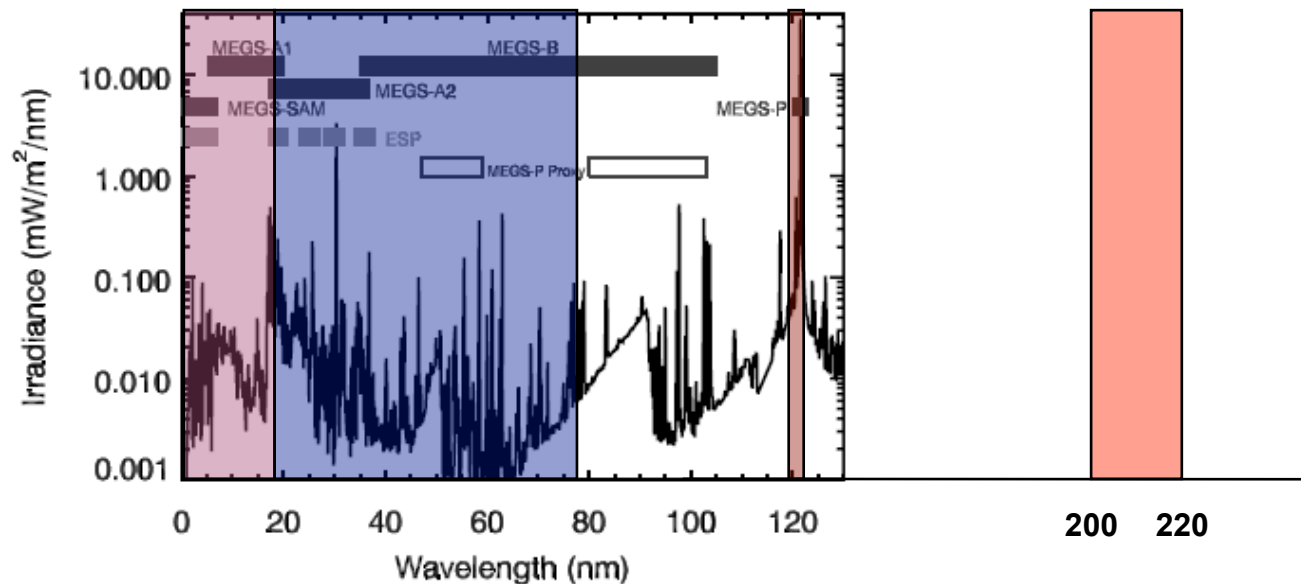
Disappearance and appearance of the active region behind the lunar disk





Cross-calibration with SDO/EVE

- ❑ EVE aging effects are tracked by flying the same instruments on-board a rocket 1x/year
- ❑ LYRA cross-calibration campaign was scheduled during the first rocket flight



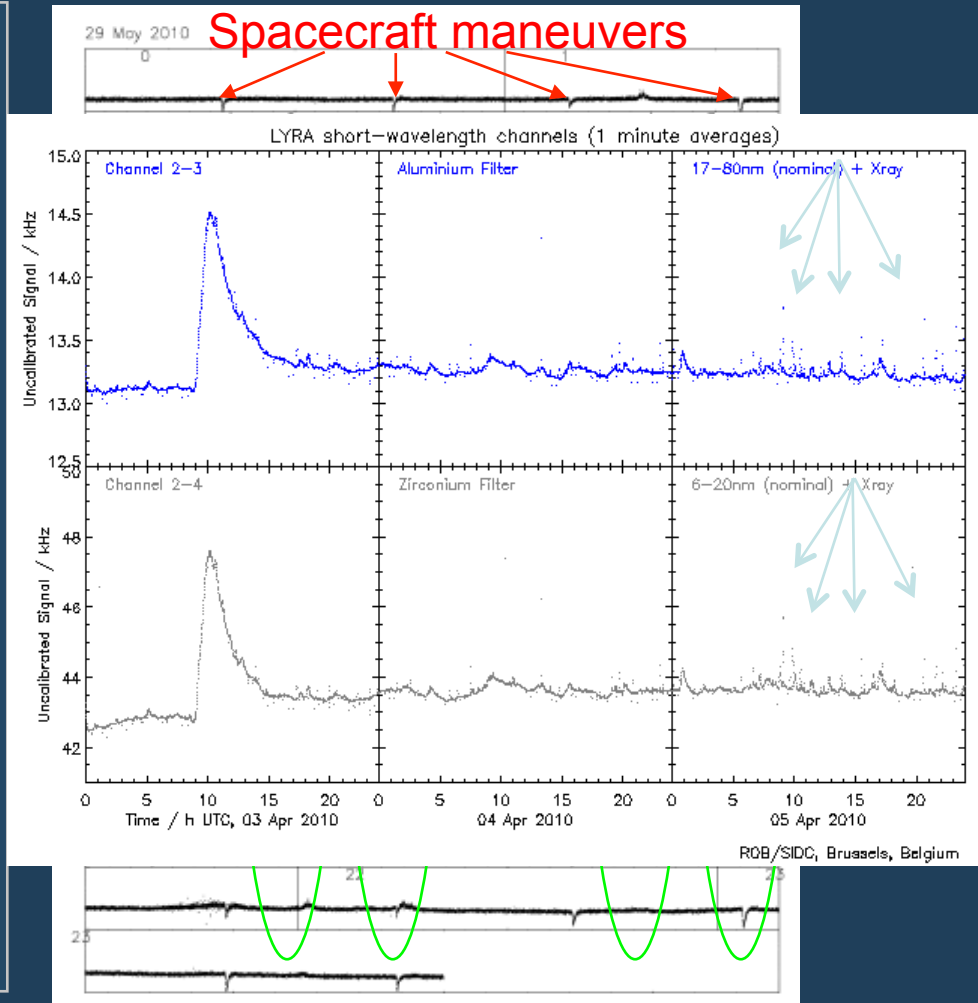
LYRA on-board PROBA2

ANNEXES



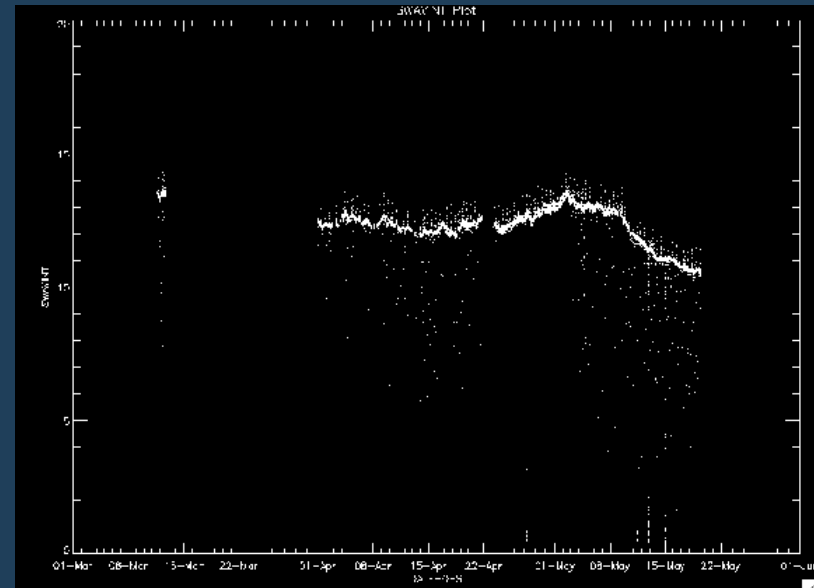
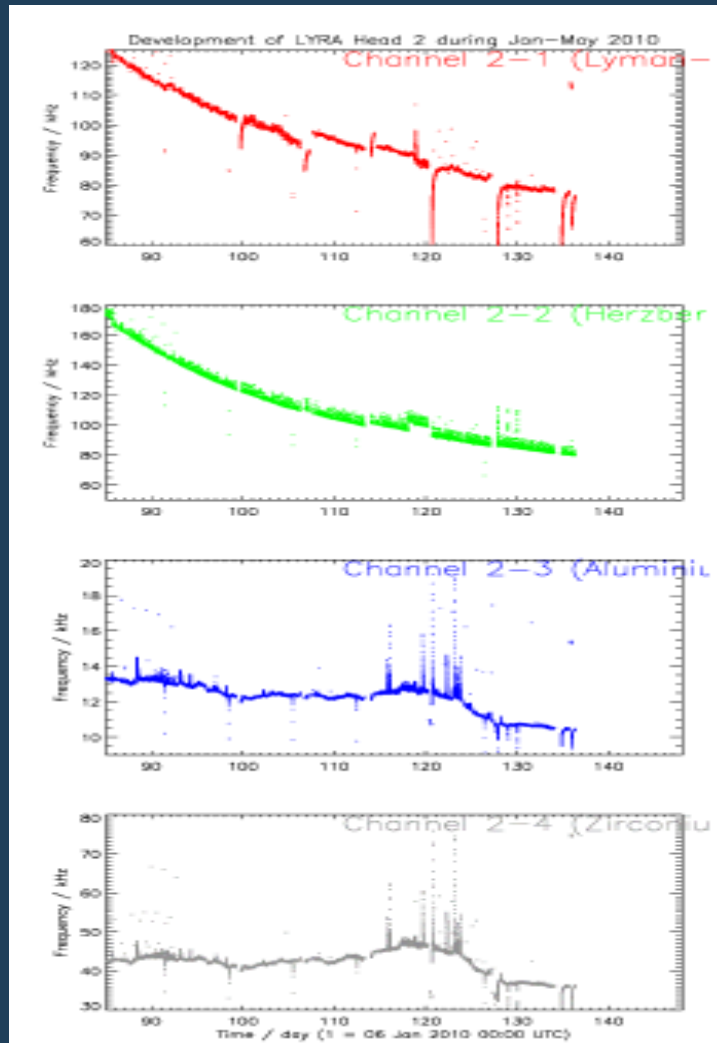
Auroral Oval

- ❑ Perturbations appearing around 75° latitude
- ❑ 2-3 days after a CME, flare ...
- ❑ Associated to geomagnetic perturbations
- ❑ Only in Al and Zr





And a fifth channel at 17.4nm...



... called SWAP!

State of the data processing pipeline

Telemetry packets

LY-TMR:

LY-EDG

LY-BSDG

Other tools



Raw data (in counts)



Lv1: Engineering data (in Hz)



Lv2: Calibrated data (in W/m^2)

Higher level
data products

Current state

Not distributed

Daily basis

Not available

Plots + flare list
available
irregular

In the future

Not distributed

After each
contact

After each
contact

systematic