P2SC-ROB-WR-725 - 20240212	P2SC Weekly report	**** ****
Period covered: Date:	'	Royal Observatory of Belgium
Written by: Approved by:	Laurence Wauters Marie Dominique	PROBA2 Science Center
То:	LYRA PI, marie.dominique@sidc.be SWAP PI, elke.dhuys@sidc.be	https://proba2.sidc.be ++ 32 (0) 2 3730559
CC:	ROB DIR, ronald@oma.be ESA Redu, Rene.Wittmann@esa.int and Marcus.De.Deus.Silva@esa.int ESA D/SRE, Joe.Zender@esa.int ESA D/TEC, Juha-Pekka.Luntama@esa.int and Melanie.Heil@esa.int	

1. Science

Solar & Space weather events

The level of solar activity¹ fluctuated between **low and high** this week.

Only M- and X-flares are mentioned, the most energetic one(s) per day are presented in **bold**:

	Monday 12 Feb	Tuesday 13 Feb	Wednesday 14 Feb	Thursday 15 Feb	Friday 16 Feb	Saturday 17 Feb	Sunday 18 Feb
Activity	moderate	low	Moderate	moderate	high	low	low
Flares	M2.6, M1.4, M1.1, M6.5		2 M1.0	M1.8	M3.0, X2.5, M1.5,	-	-

¹ See appendix. All timings are given in UT.

Solar Activity

Solar flare activity fluctuated from low to high during the week.

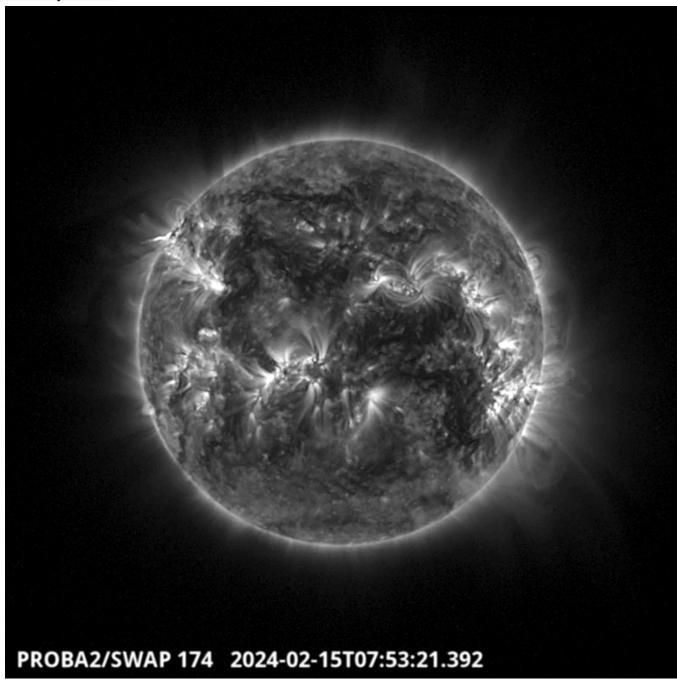
In order to view the activity of this week in more detail, we suggest to go to the following website from which all the daily (normal and difference) movies can be accessed: https://proba2.oma.be/ssa
This page also lists the recorded flaring events.

A weekly overview movie can be found here (SWAP week 725).

Details about some of this week's events can be found further below.

If any of the linked movies are unavailable they can be found in the P2SC movie repository here

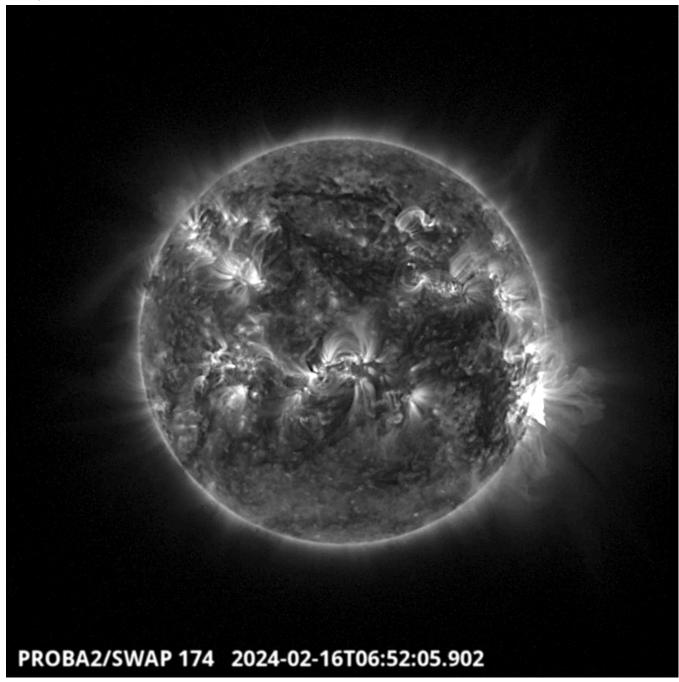
Thursday Feb 15



A prominence eruption has been produced around 7:53 UR on the North East limb of the solar disk. It has been followed by a C5.1 flare (AR 3586).

Find a SWAP movie of the event here.

Friday Feb 16



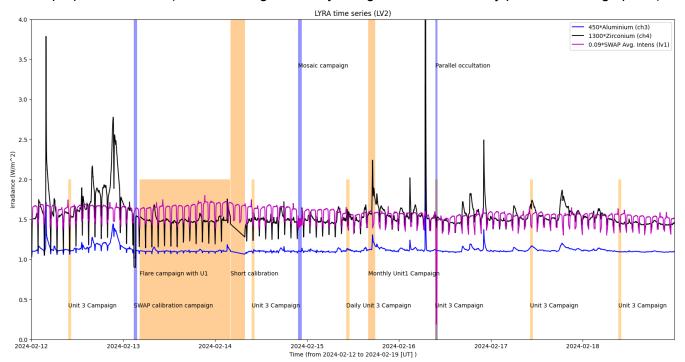
The strongest flare of this week is a X2.5, it is visible on the SWAP image above around 6:52 UT. It has been produced by the AR NOAA 3576, located in the South West part of the solar disk.

Find a SWAP movie of the event here.

An overview of the weekly LYRA & SWAP data is provided below:

The following curves are visible:

- black: Zirconium Channel LYRA Unit 2
- blue: Aluminium Channel of LYRA Unit 2
- purple: SWAVINT (SWAP Average Intensity; integrated solar intensity per SWAP image pixel)



Operations and Calibrations:

The blue shaded periods related to SWAP, correspond to, from left to right:

- SWAP calibration, 2024-Feb-13
- Mosaïc campaign, 2024-Feb-14
- SWAP parallel occultation campaign with LYRA, 2024-Feb-16

The orange shaded periods related to LYRA correspond to, from left to right:

- Daily unit 3 campaign during occultation.
- LYRA Unit1 flare campaign, 2024-Feb-13
- Short calibration, 2024-Feb-14
- Monthly unit1 campaign, 2024-Feb-15

The red shaded periods related to other issues corresponds to:

None

2. LYRA instrument status

IOS

Start IOS	Mon Feb 12 2024	LYIOS01055
End IOS	Sun Feb 18 2024	LYIOS01056

LYRA detector temperature

LYRA detector 2 temperature globally varied between 52.49 and 56.73 $^{\circ}\text{C}.$

3. SWAP instrument status

MCPM errors

The number of MCPM recoverable errors increased from 52511 to 52963.

The number of MCPM unrecoverable errors remained at 3135.

IOS

Start IOS	Mon Feb 12 2024	IOS01181
End IOS	Sun Feb 18 2024	IOS01182

SWAP detector temperature

The SWAP Cold Finger Temperature globally varied between 2.47 and 4.07 °C.

4. PROBA2 Science Center Status

The following changes were made to the P2SC:

• Several configuration parameters have been changed to avoid numerous warnings occurring at the end of this occultation period.

5. Data reception & discussions with MOC

Passes

The delivery of the passes for this week (passes 46389 to 46447) was nominal, except for:

None.

Data coverage HK

All HK data files (LYRA_AD) have been received, except:

None.

Data coverage SWAP

All SWAP Science data files (BINSWAP) have been received, except:

None.

Total number of images between 2024 Feb 12 00:00 UT and 2024 Feb 19 00:00 UT: 4056

Highest cadence in this period: 18 seconds Average cadence in this period: 148.87 seconds Number of image gaps larger than 300 seconds: 150

Largest data gap: 21.47 minutes

Data coverage LYRA

All LYRA Science data files (BINLYRA) have been received, except:

None

6. APPENDIX: Frequently used acronyms

ADPMS Advanced Data and Power Management System

AOCS Attitude and Orbit Control System

APS Active Pixel image Sensor

ASIC Application Specific Integrated Circuit

BBE Base Band Equipment
CME Coronal Mass Ejection

COGEX Cool Gas Generator Experiment

CRC Cyclic Redundancy Check
DAC Data Acquisition Controller
DBR Deployment, backup & recovery
DDA Decommutated data archive
ESP Experimental Solar Panel

FITS Flexible Image Transport System

FOV Field Of View FPA Focal Plane Assembly

FPGA Field Programmable Gate Arrays

GPS Global Positioning System

HK Housekeeping

IOS Instrument Operations Sheet

LED Light Emitting Diode
LYRA LYman alpha RAdiometer

LYTMR LYRA Telemetry Reformatter (software module of P2SC)
LYEDG LYRA Engineering Data Generator (software module of P2SC)

MCPM Mass Memory, Compression and Packetisation Module

MOC Mission Operation Center NDR Non Destructive Readout

OBSW On board Software
PI Principal Investigator
P2SC PROBA2 Science Center
ROB Royal Observatory of Belgium

SAA South Atlantic Anomaly
SEU Single Event Upset

SoFAST | Solar Feature Automated Search Tool

SWAP Sun Watcher using APS detector and image Processing

SWAVINT | SWAP AVerage INTensity

SWBSDG | SWAP Base Science Data Generator

SWEDG SWAP Engineering Data Generator (software module of P2SC)
SWTMR SWAP Telemetry Reformatter (software module of P2SC)

TBC To Be Confirmed

TBD To Be Defined TC Telecommand

UTC Coordinated Universal Time

UV Ultraviolet

VFC Voltage to Frequency Converter

7. APPENDIX Solar Activity Definitions

In the science section we use the following solar activity standards.

The standard scale for solar activity is:

- very low (almost no flares, only B)
- low (a few C flares)
- moderate (many C flares and at least an M flare)
- high (several M flares and an X flare)
- very high (continuous background of C flares, numerous M flares, more than one X flare)