SVAP & LYRA: READY FOR SCIENCE & SPACE WEATHER

Anik De Groof & PROBA2 Science Center team ESA c/o Royal Observatory of Belgium

CESRA 2010 meeting 🔆 La Roche en Ardenne, Belgium 🔆 June 16, 2010







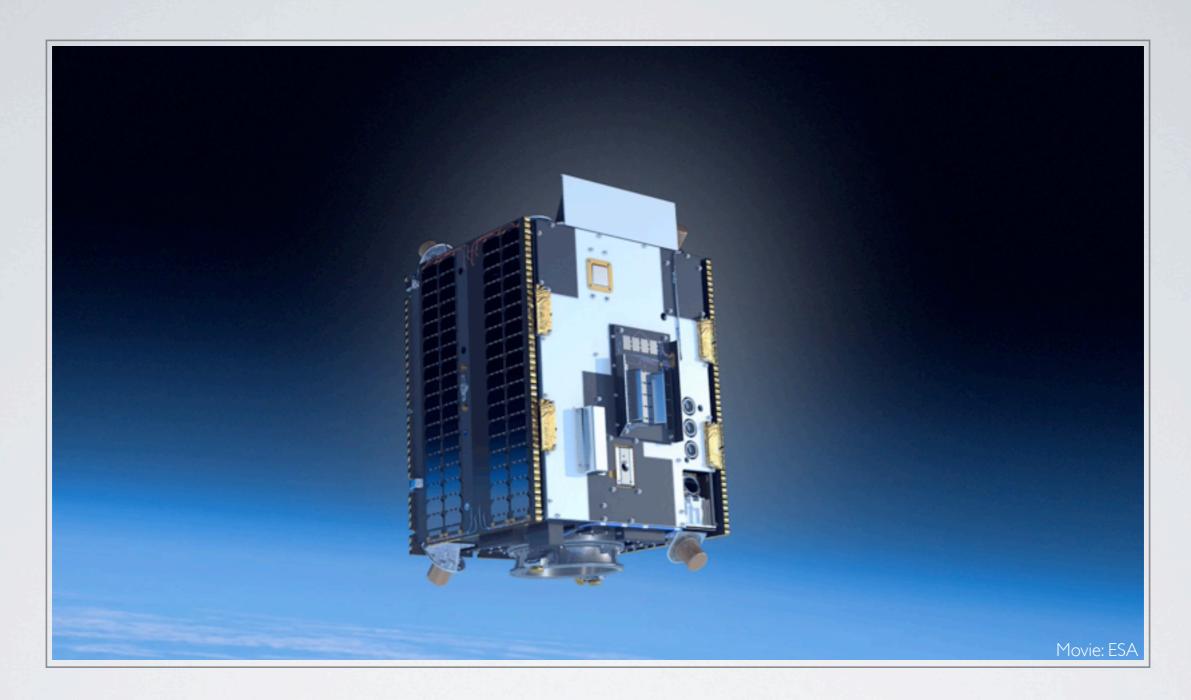






PROBA2

Project for On-Board Autonomy
Microsatellite in sun-synchronous orbit 725 km altitude
Launched on Nov. 2, 2009



ESATECHNOLOGY MISSION

4 innovative instruments: SWAP, LYRA, TPMU, DSLP 17 technological experiments in-orbit demonstration



ESA SCIENCE MISSION

SWAP and LYRA observe the Sun in EUV and XUV nominal operations since March



PROBA2 SCIENCE CENTER

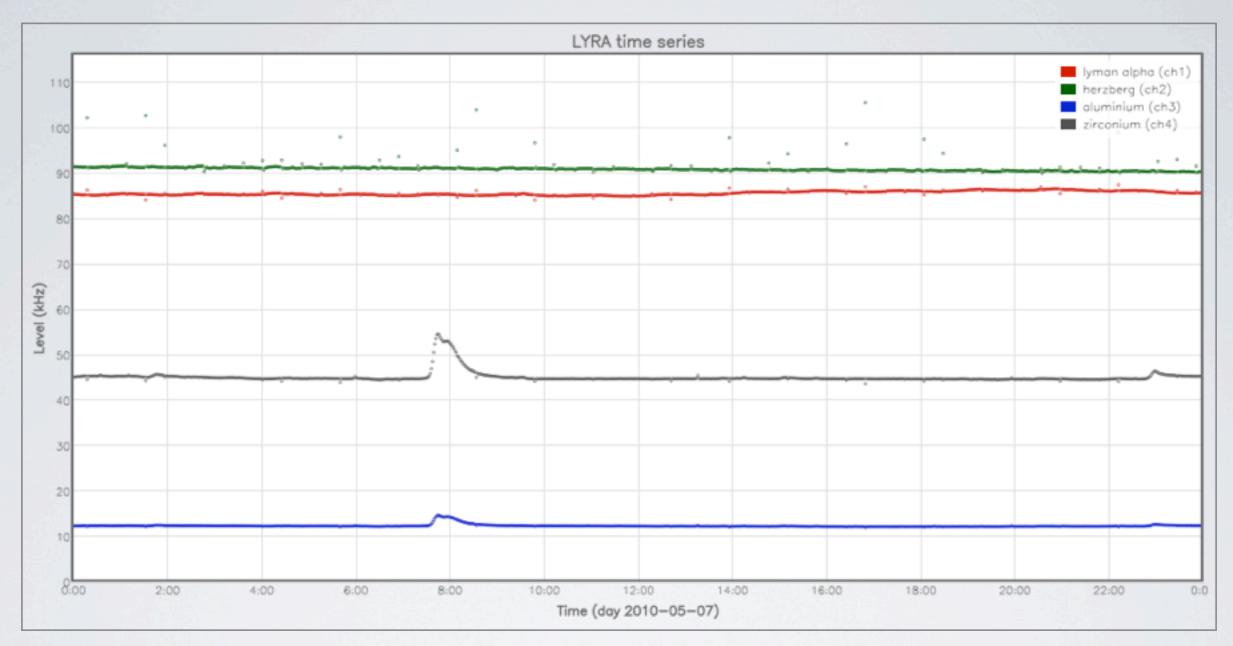
Instrument commanding + data processing & storage at ROB Data available in near-real time (~Ihr after contact)

LYRA: LYMAN-ALPHA RADIOMETER

- 3 instrument units (redundancy)
- 4 spectral channels per head
- 3 types of detectors: Silicon &
 - 2 types diamond detectors:
 - radiation resistant
 - insensitive to visible light compared to Si detectors
- High cadence up to 100Hz



LYRA CHANNELS

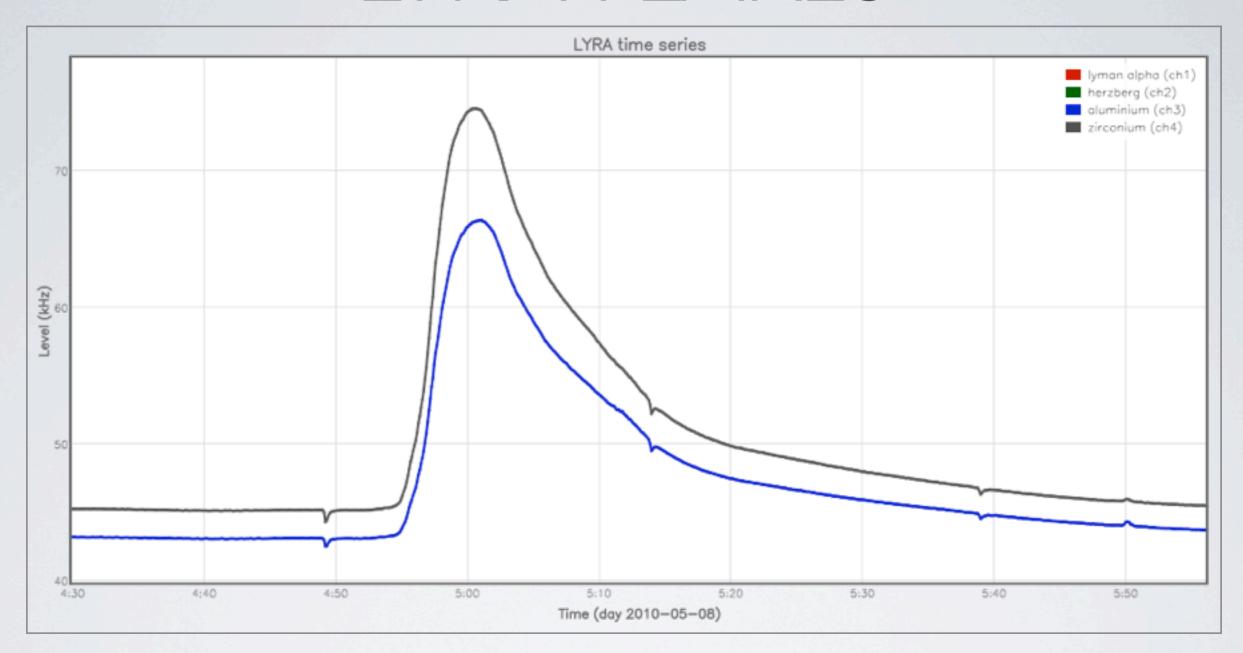


Lyman-α: 120-123 nm

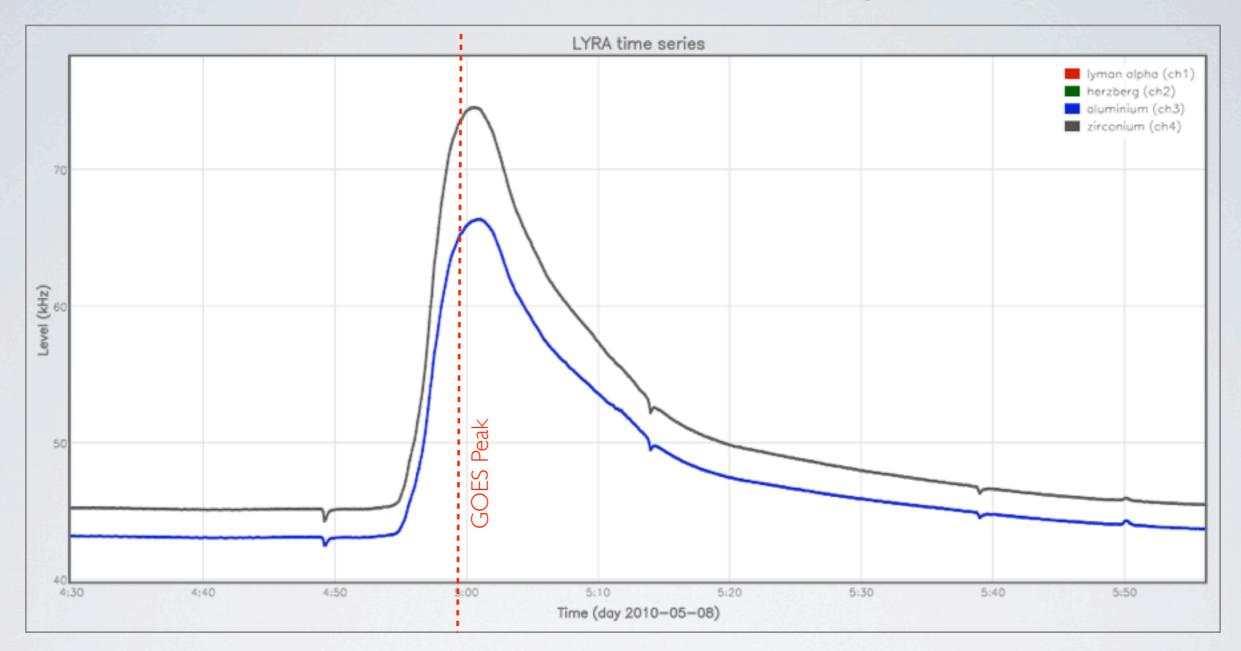
Herzberg Continuum: 200-220 nm

Aluminum Filter: XUV & EUV (incl. He II at 30.4 nm) ~ 17-80 nm

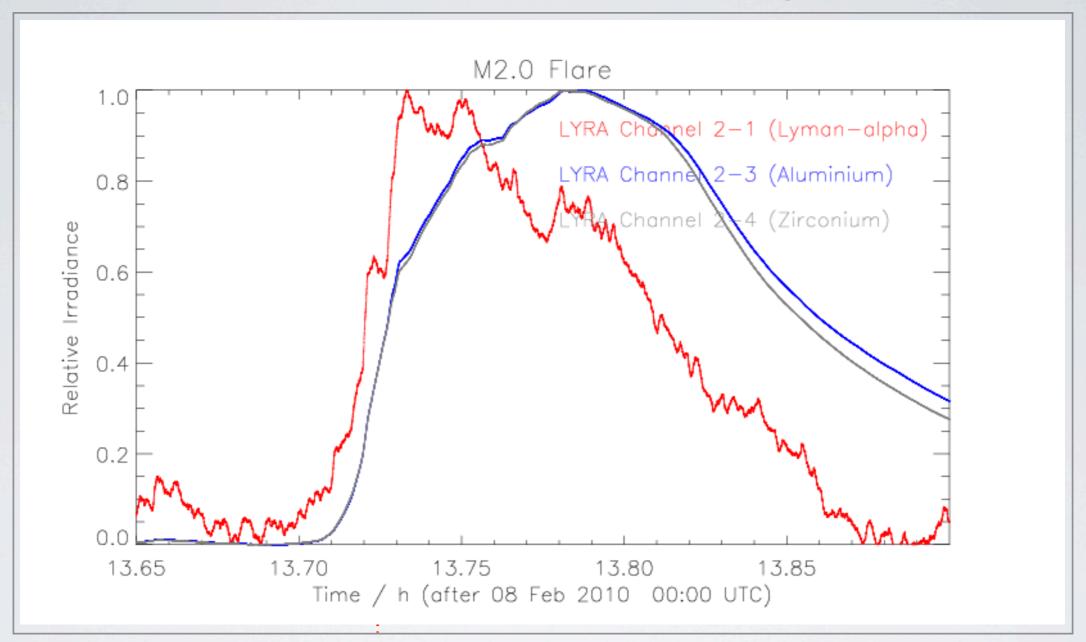
Zirconium Filter: XUV & EUV (excl. He II) ~ 1-20nm



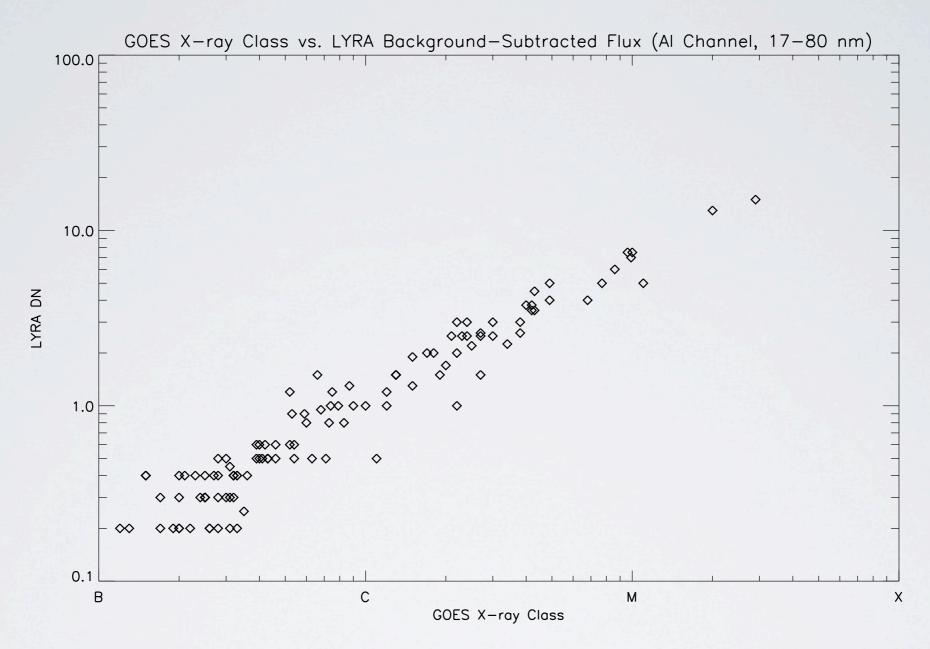
LYRA senses all flares in Zr & Al up to 10ms resolution
Ly-α contribution for impulsive flares
Different onset & peak times in different pass bands
Good correlation to GOES flares with better temporal resolution



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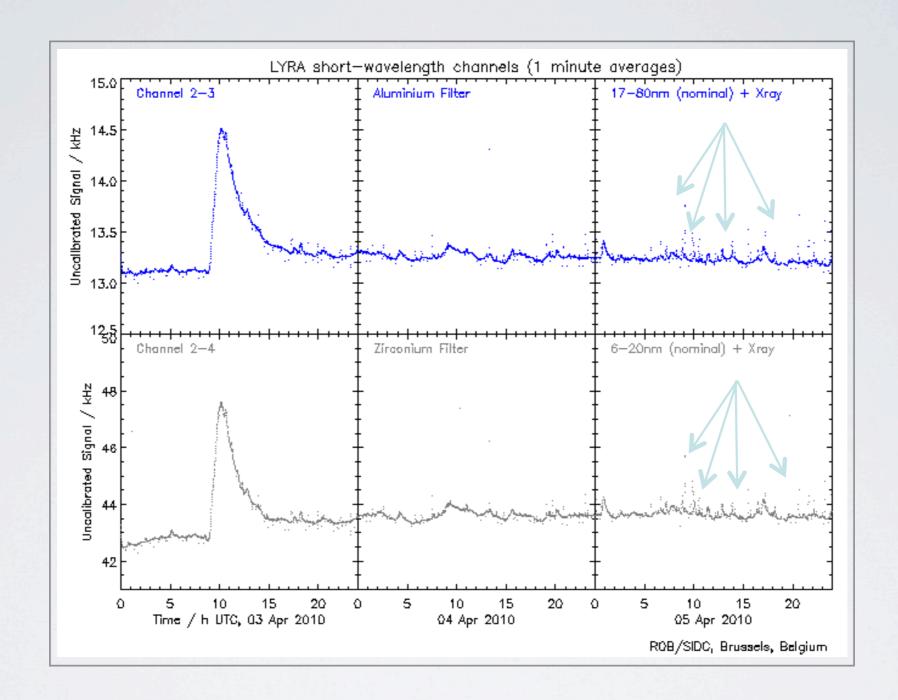
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Ly-\alpha contribution for impulsive flares

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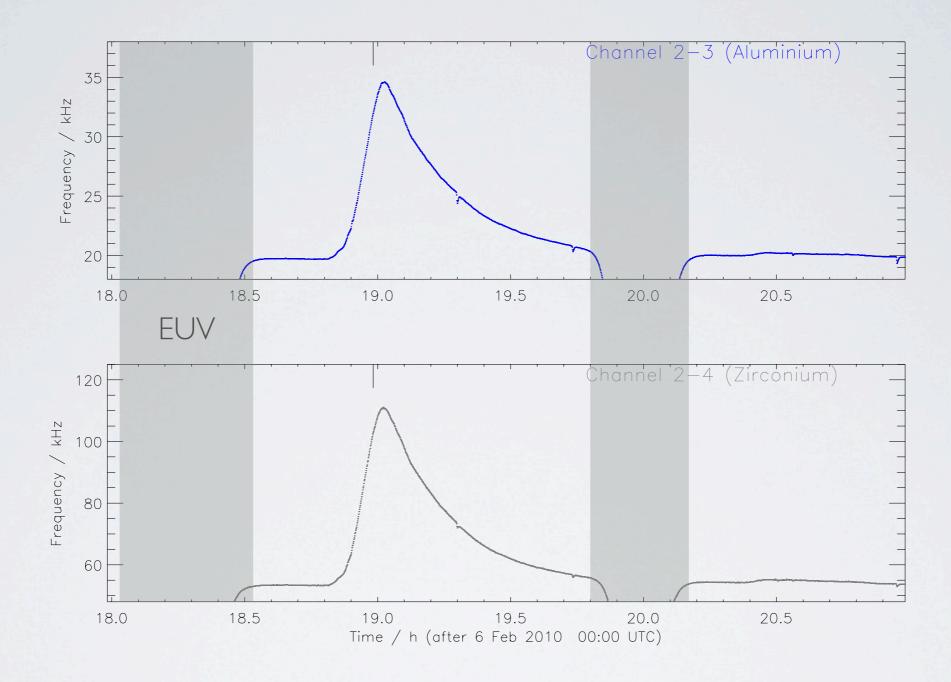
Good correlation to GOES flares with better temporal resolution

OTHER LYRA EVENTS



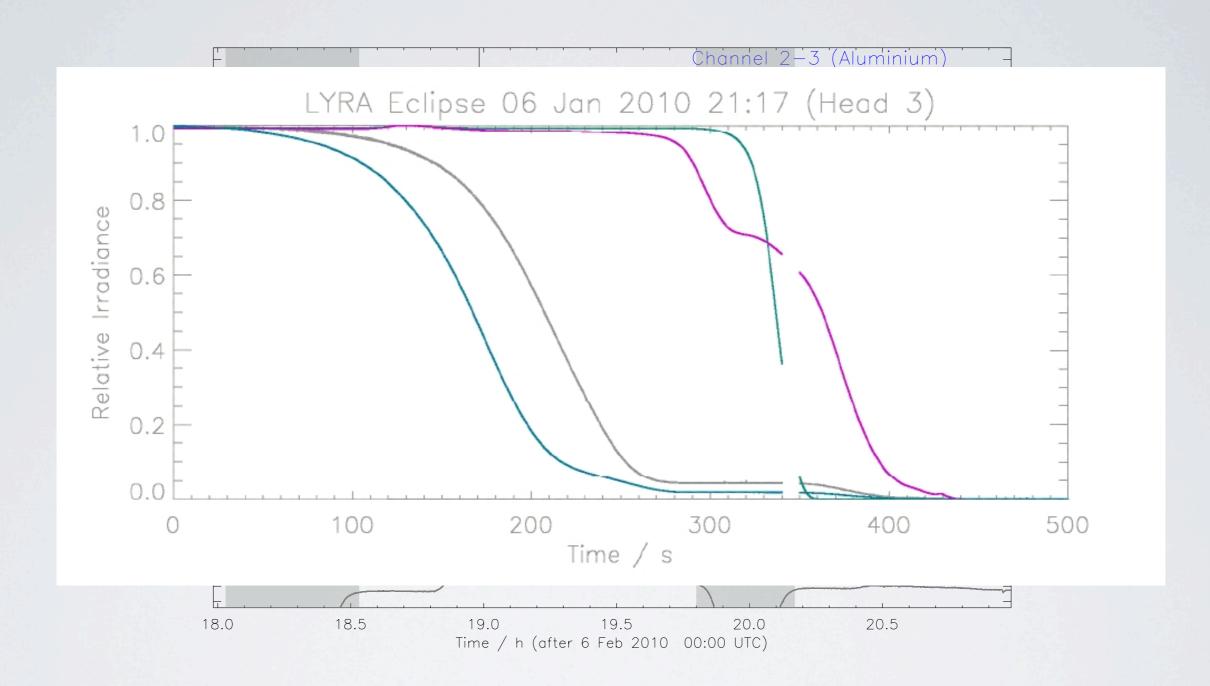
Geomagnetic perturbations: around +-75deg latitude, 2-3 days after a CME

OTHER LYRA EVENTS

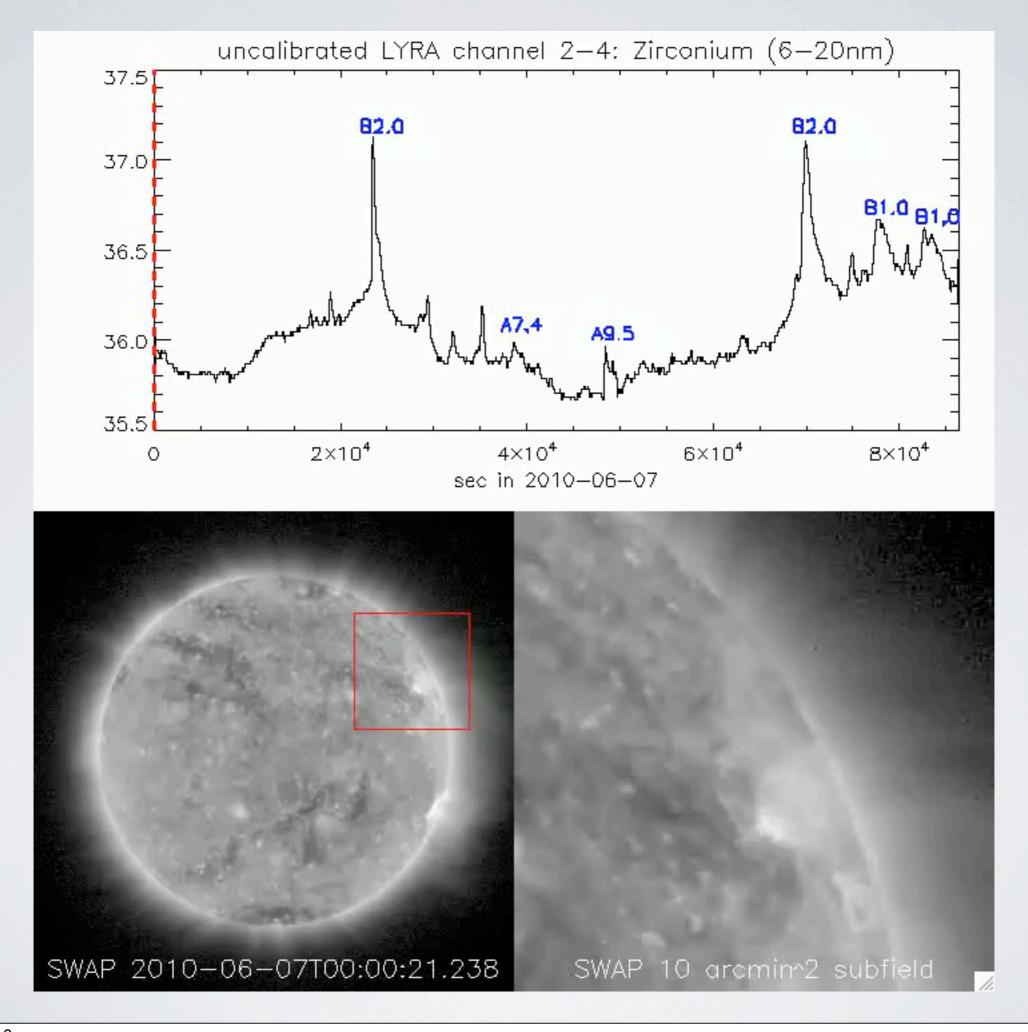


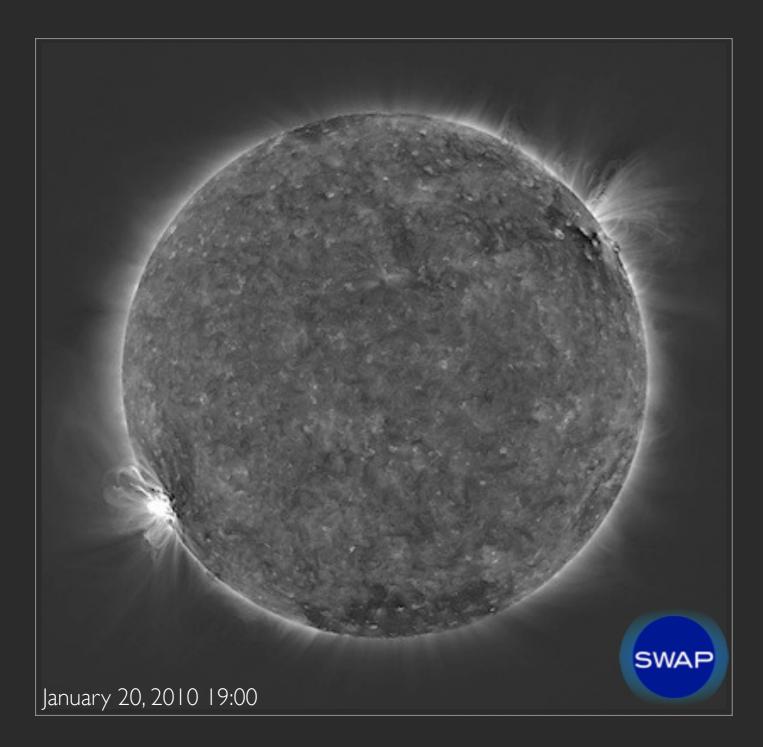
Nov-Feb: eclipse occultations by Earth's atmosphere & Earth

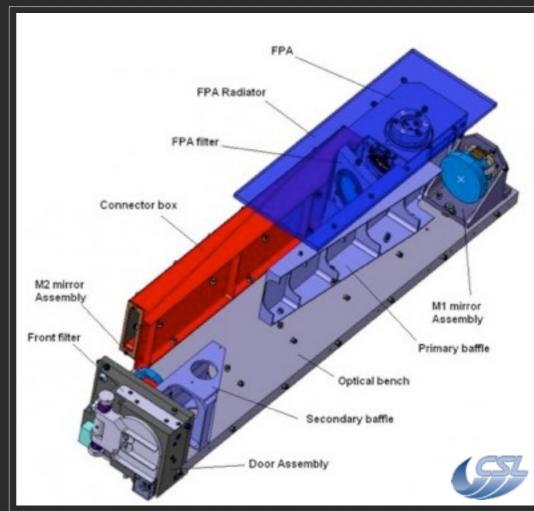
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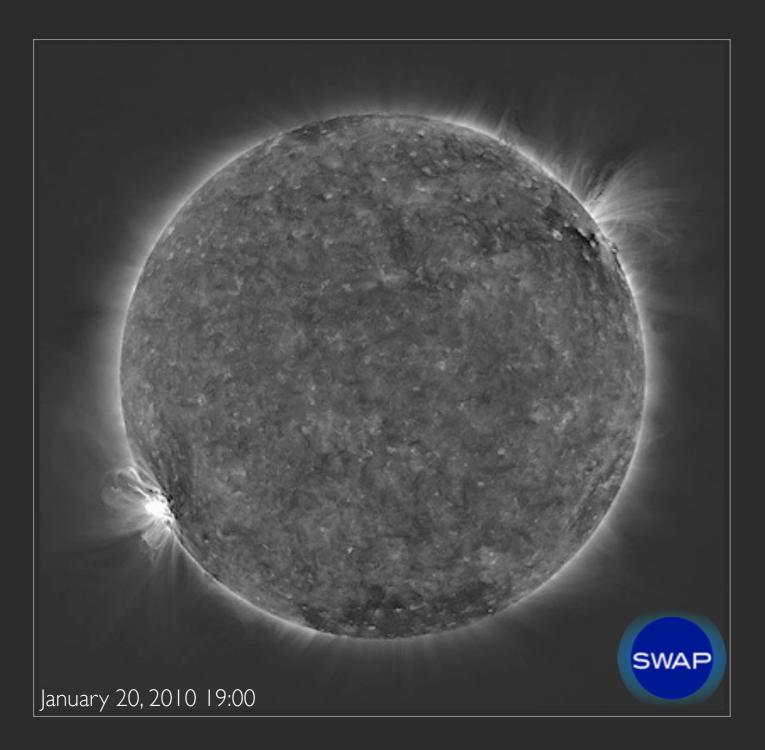


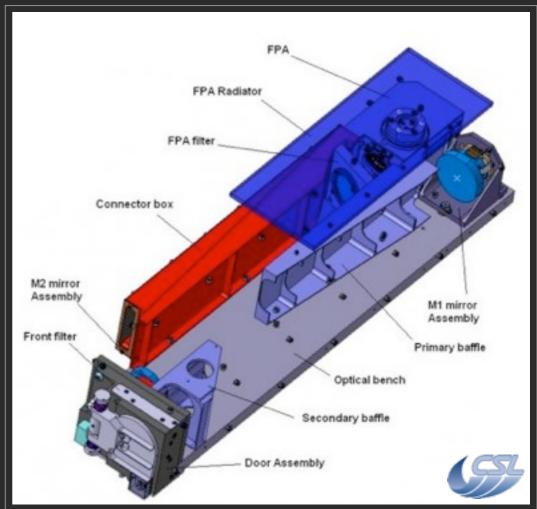




SWAP EUV IMAGER

Observes the I million degree corona in EUV light



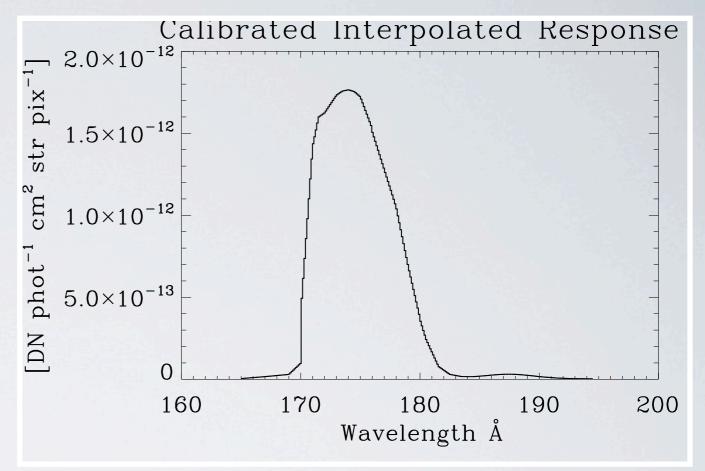


SWAP EUV IMAGER

Observes the 1 million degree corona in EUV light Exercise in miniaturization: off-axis Ritchey-Chrétien scheme

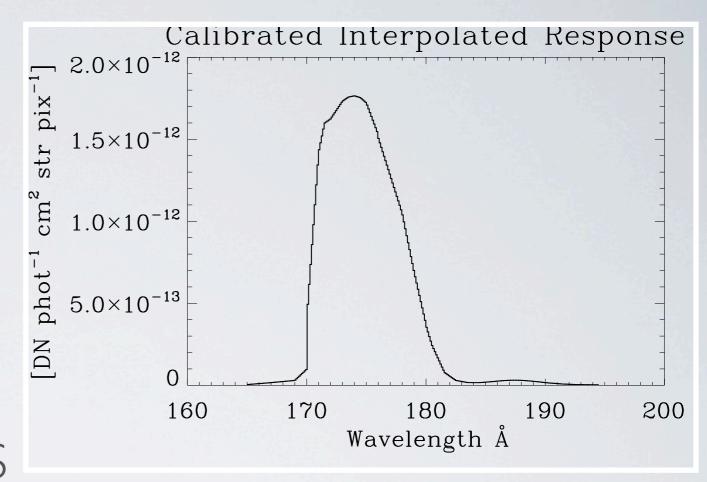
SPECTRAL RESPONSE

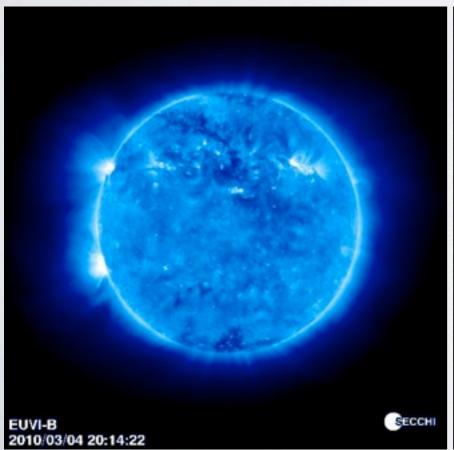
Peak at 17.4nm
Comparable to EIT 17.1nm
& STEREO/EUVI 17.1nm
~Imin cadence, IK xIK CMOS

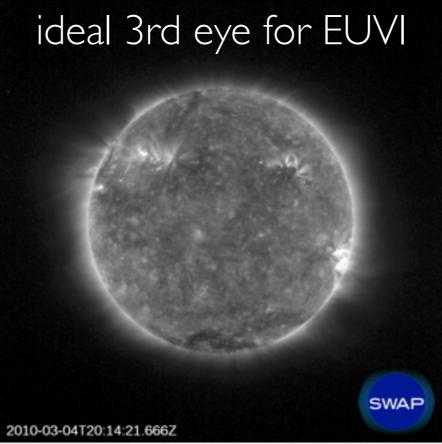


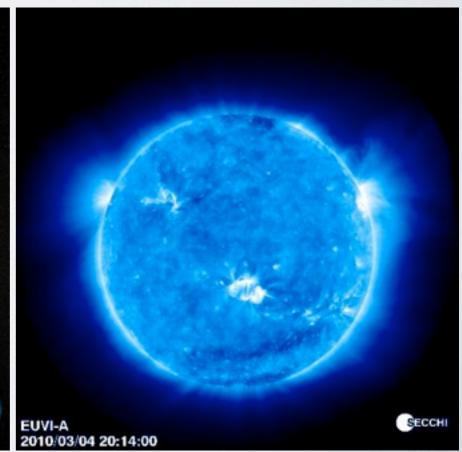
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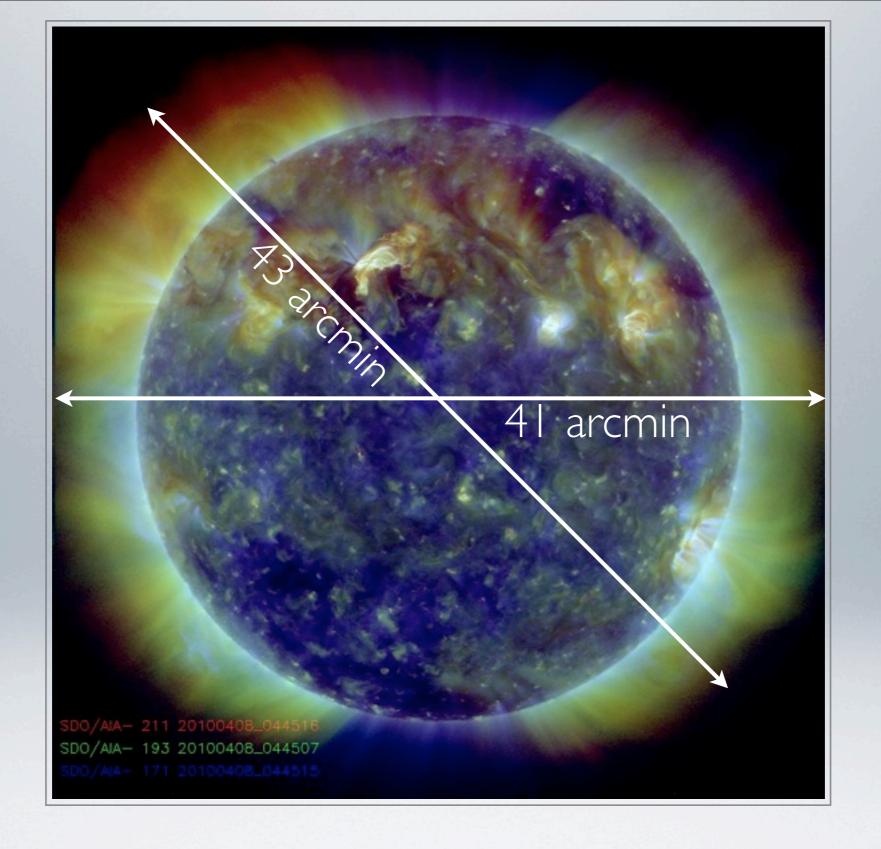
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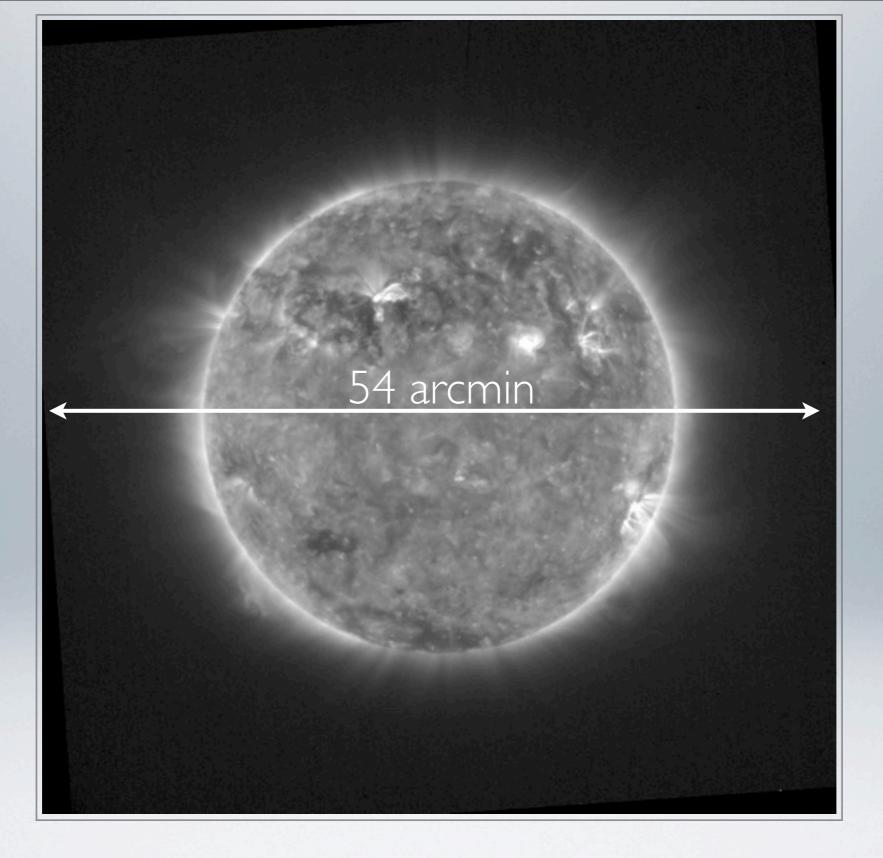




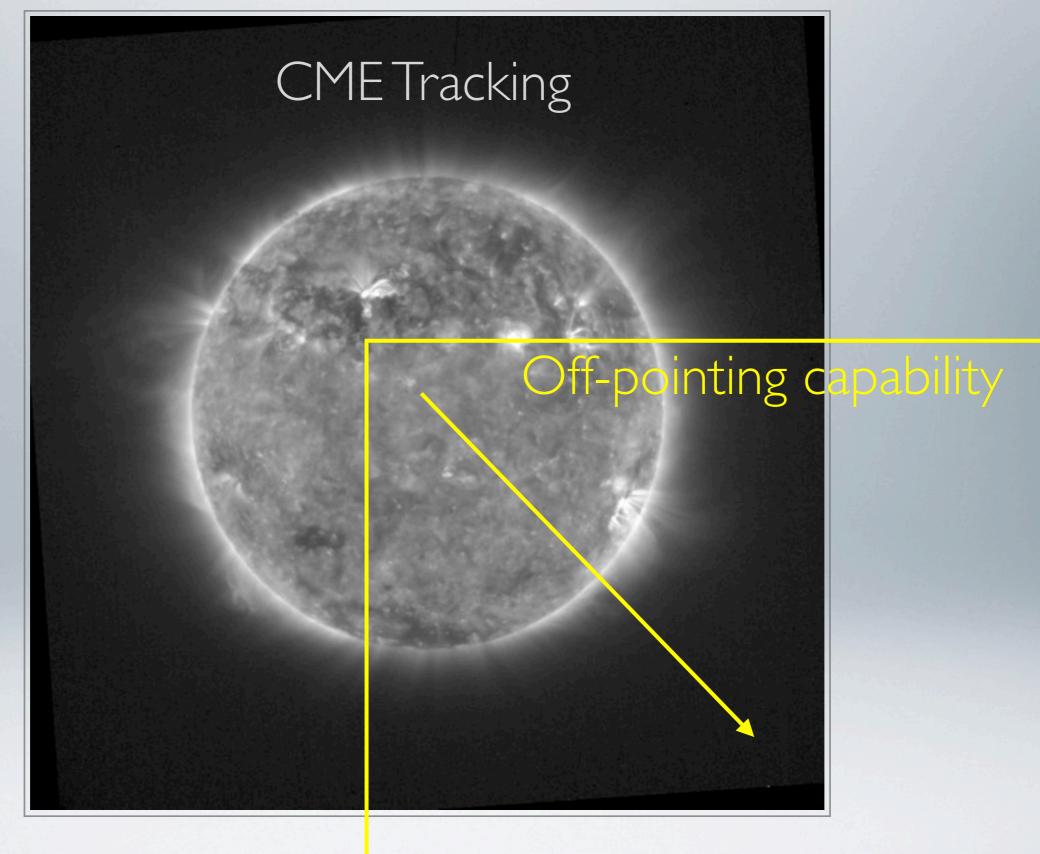




SWAP VS AIA ON SDO



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ANNULAR SOLAR ECLIPSE



January 15, 2010, 06:00 UTC

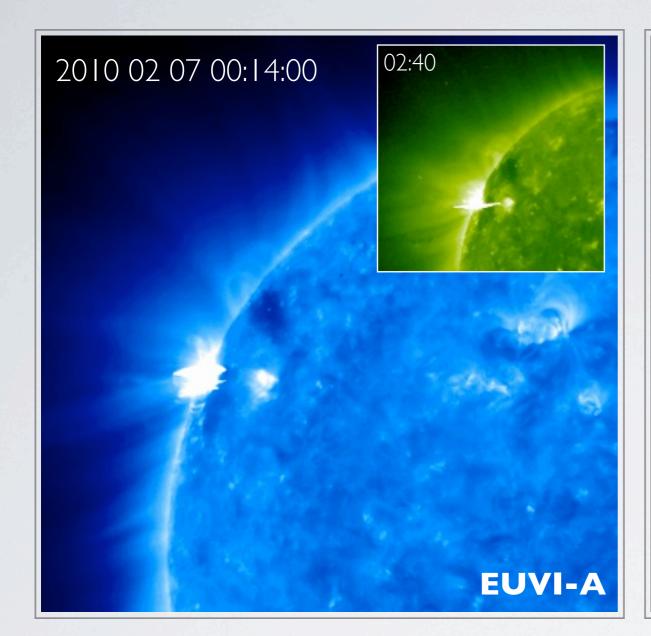


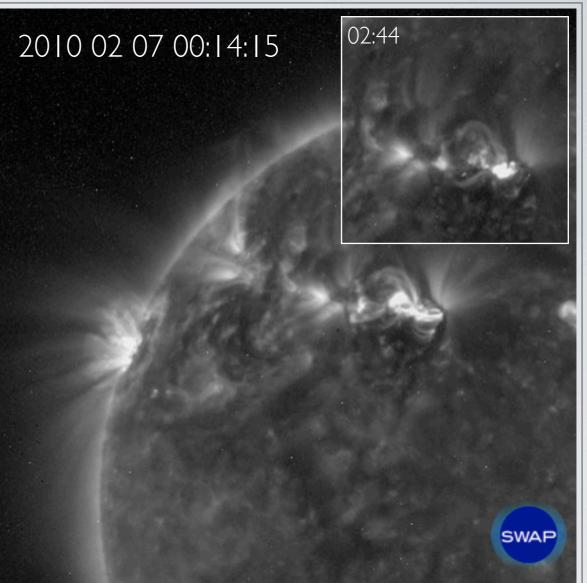
ACTIVE REGIONS & FLARES



January 18-21, 2010

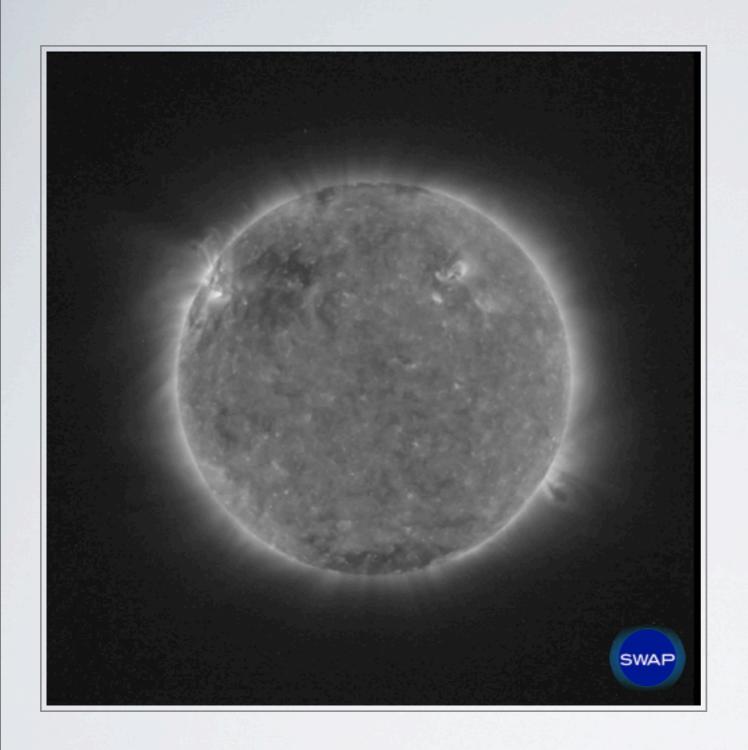


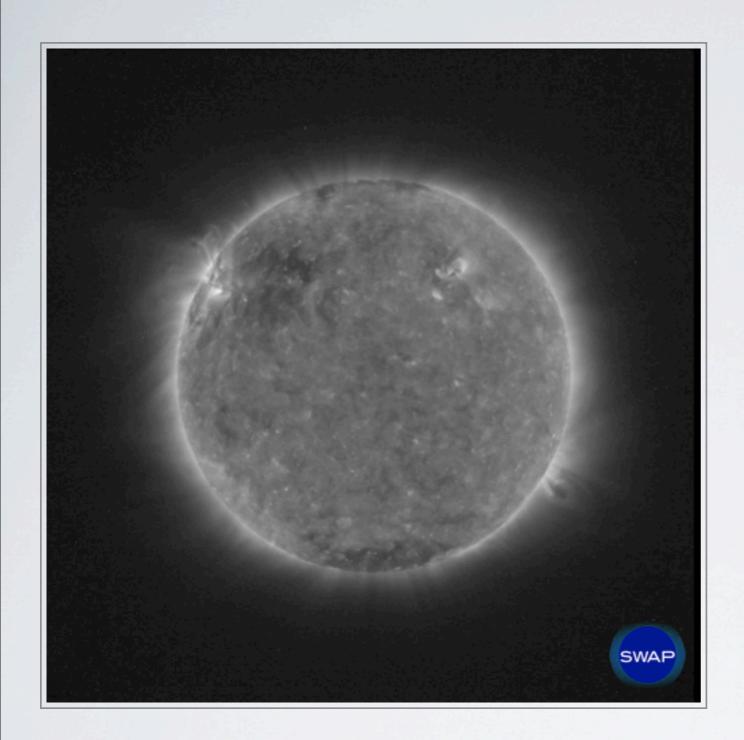




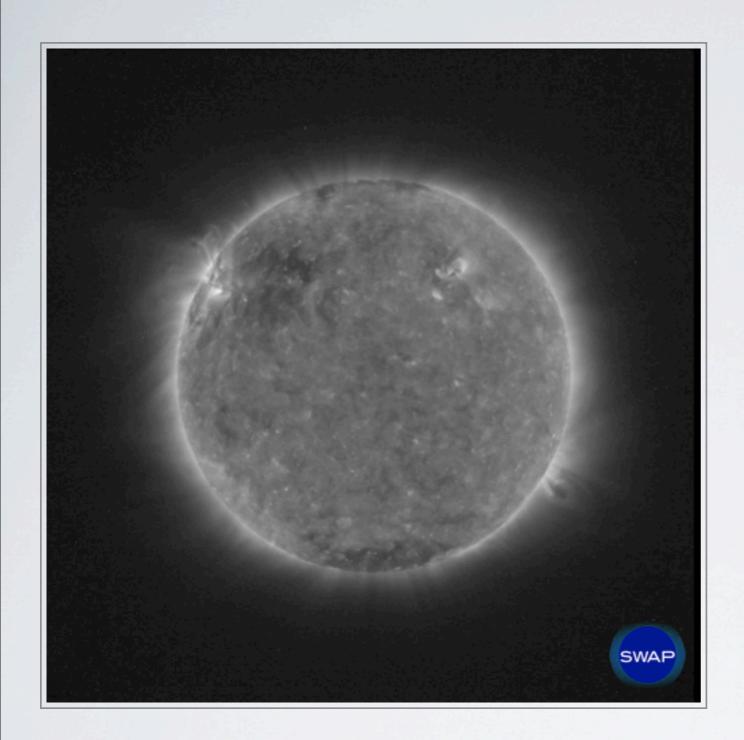
SOLAR FLARES

limited blooming due to CMOS detector high cadence up to 18s LYRA gives detailed temporal evolution (10ms)

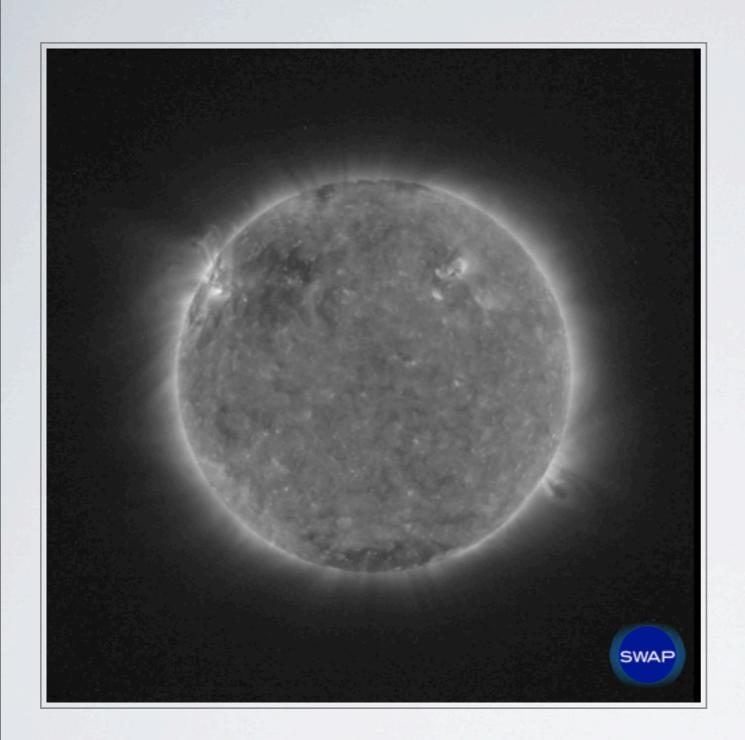


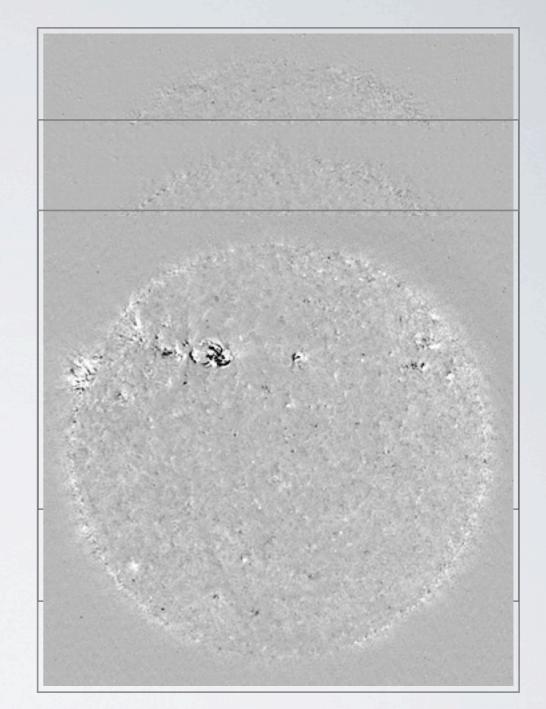


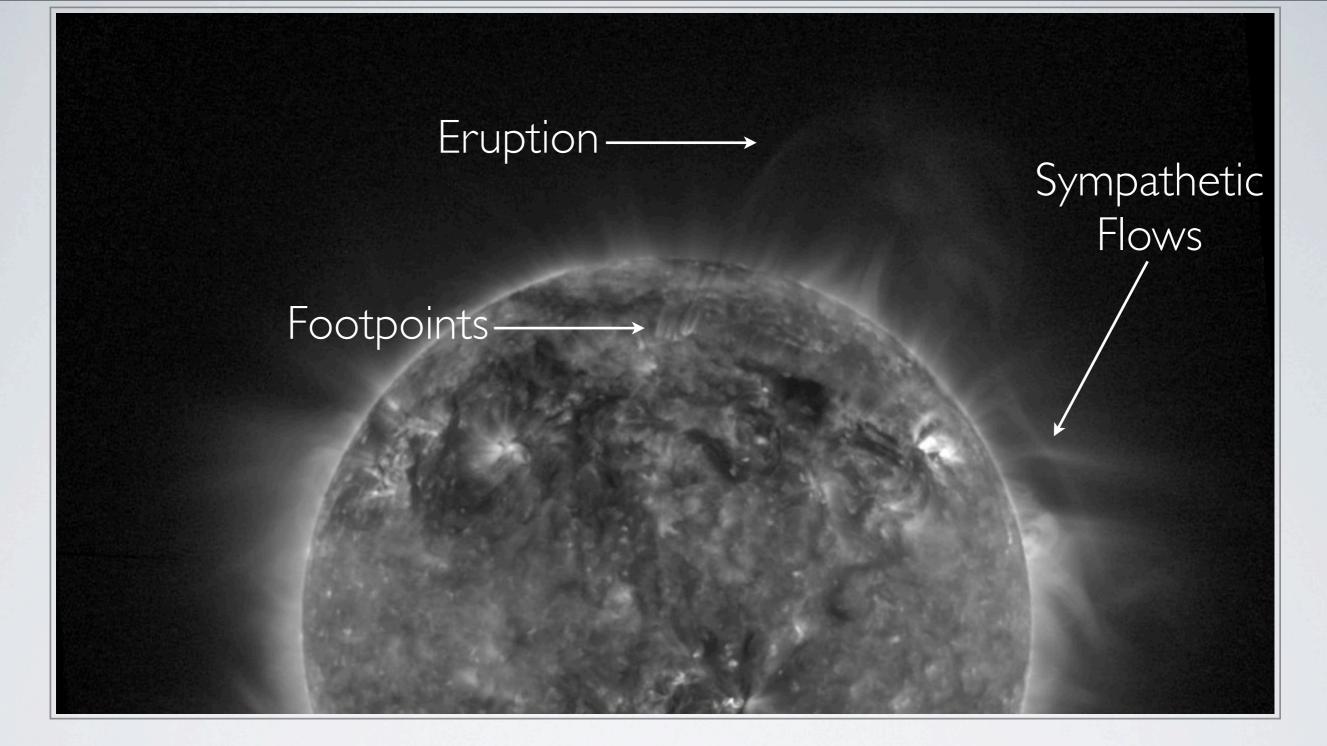






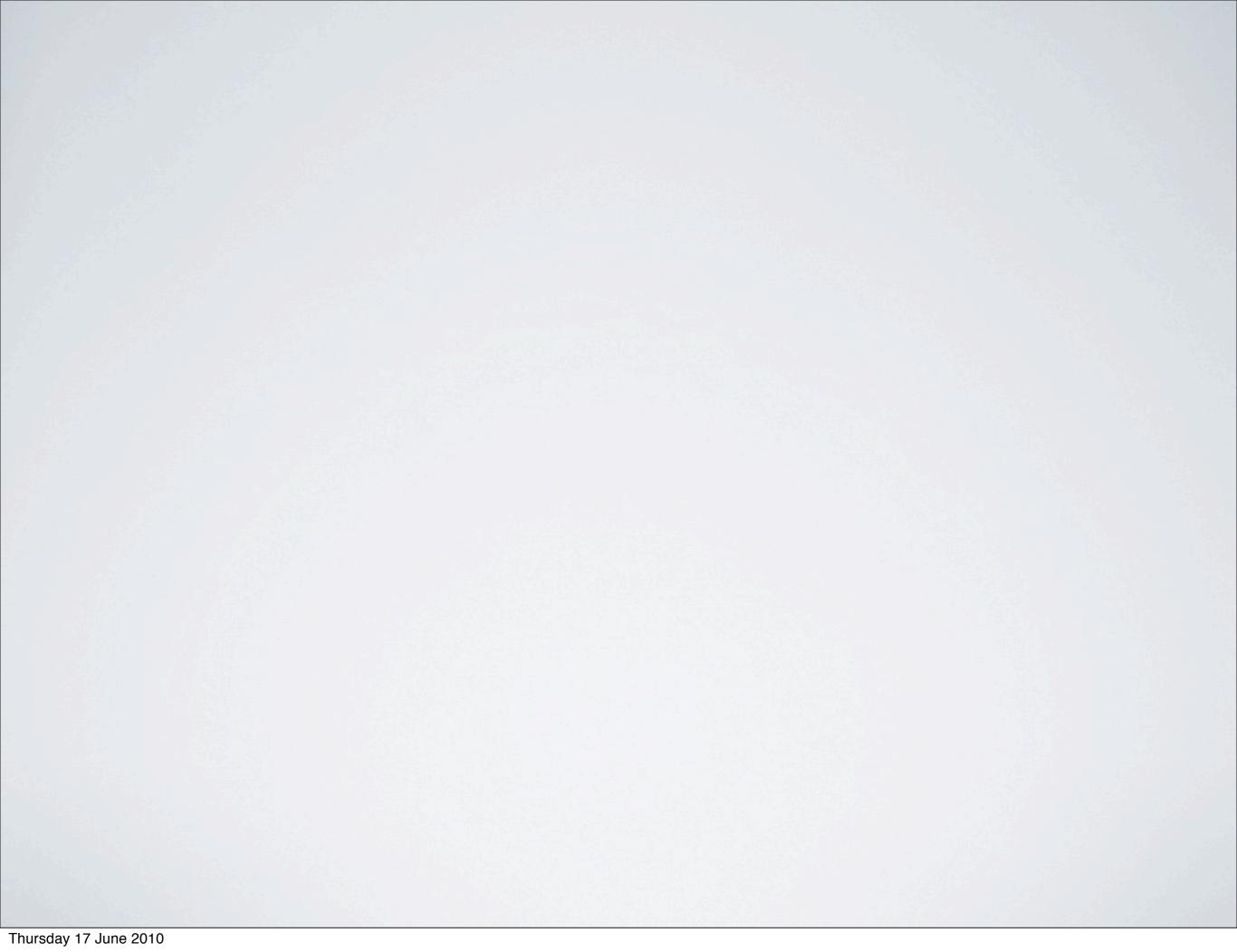


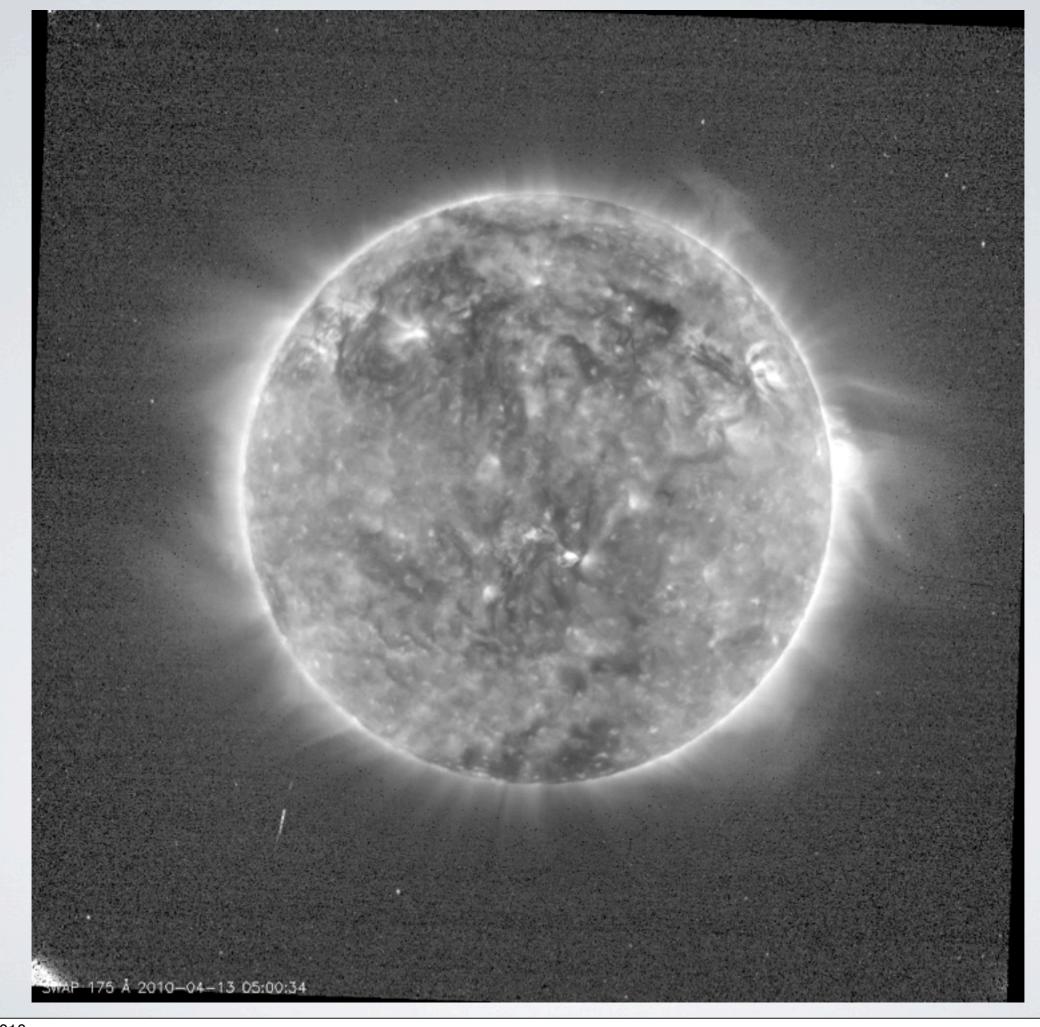


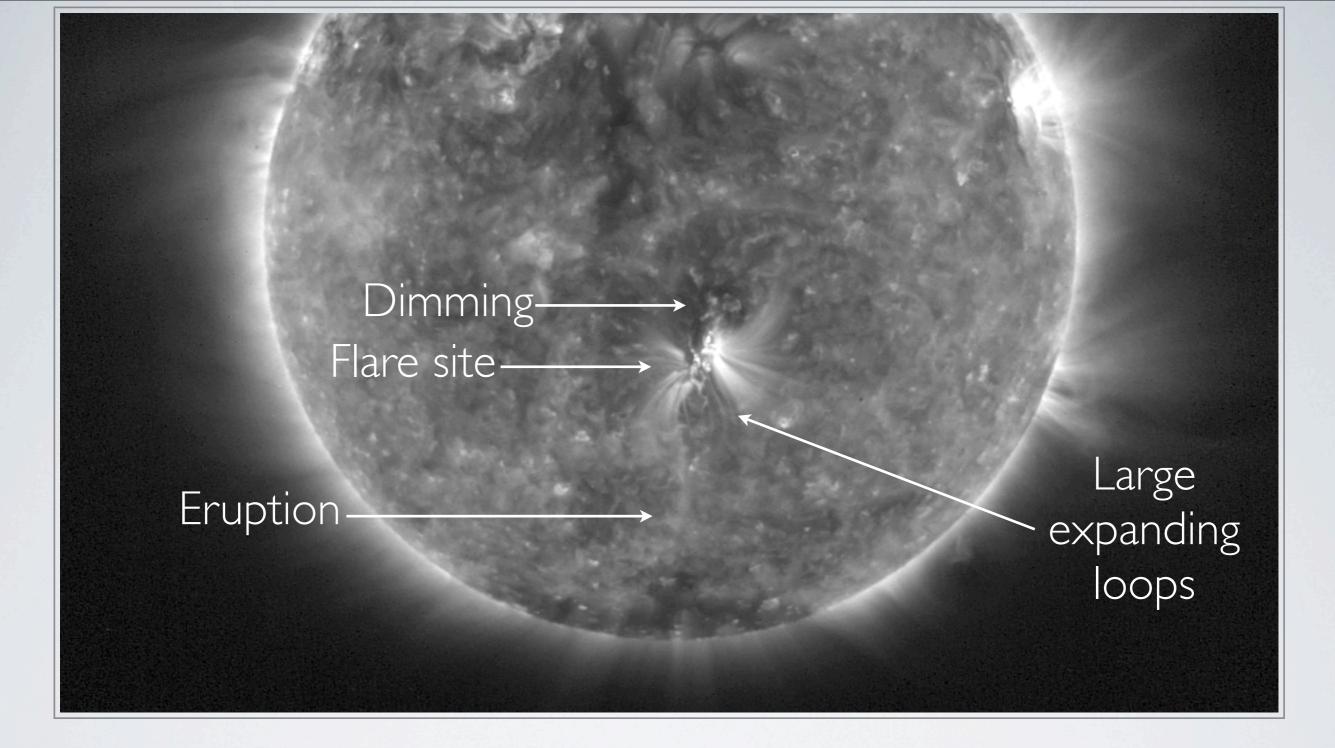


PROMINENCE ERUPTION

13 April 2010, 09:30 UTC

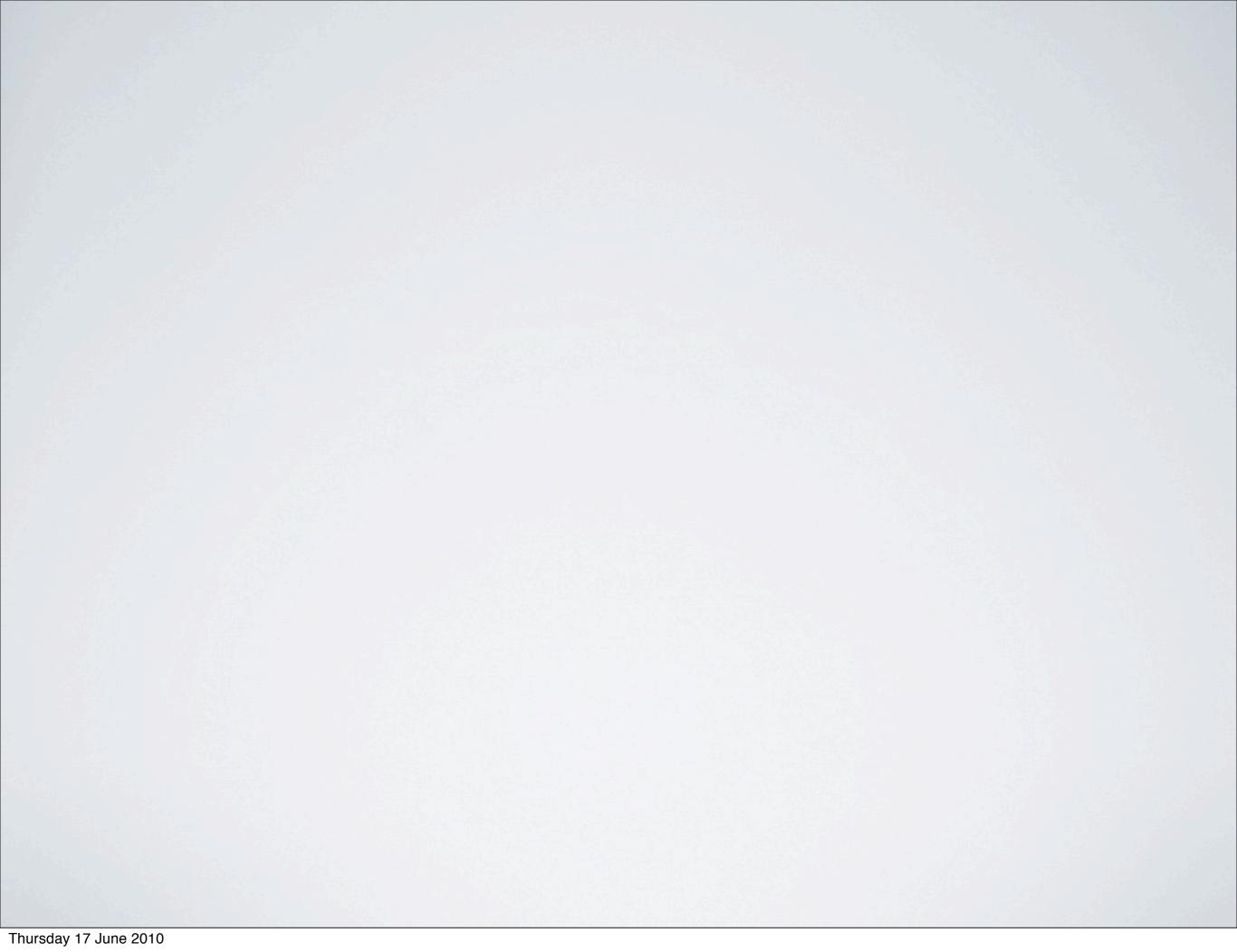




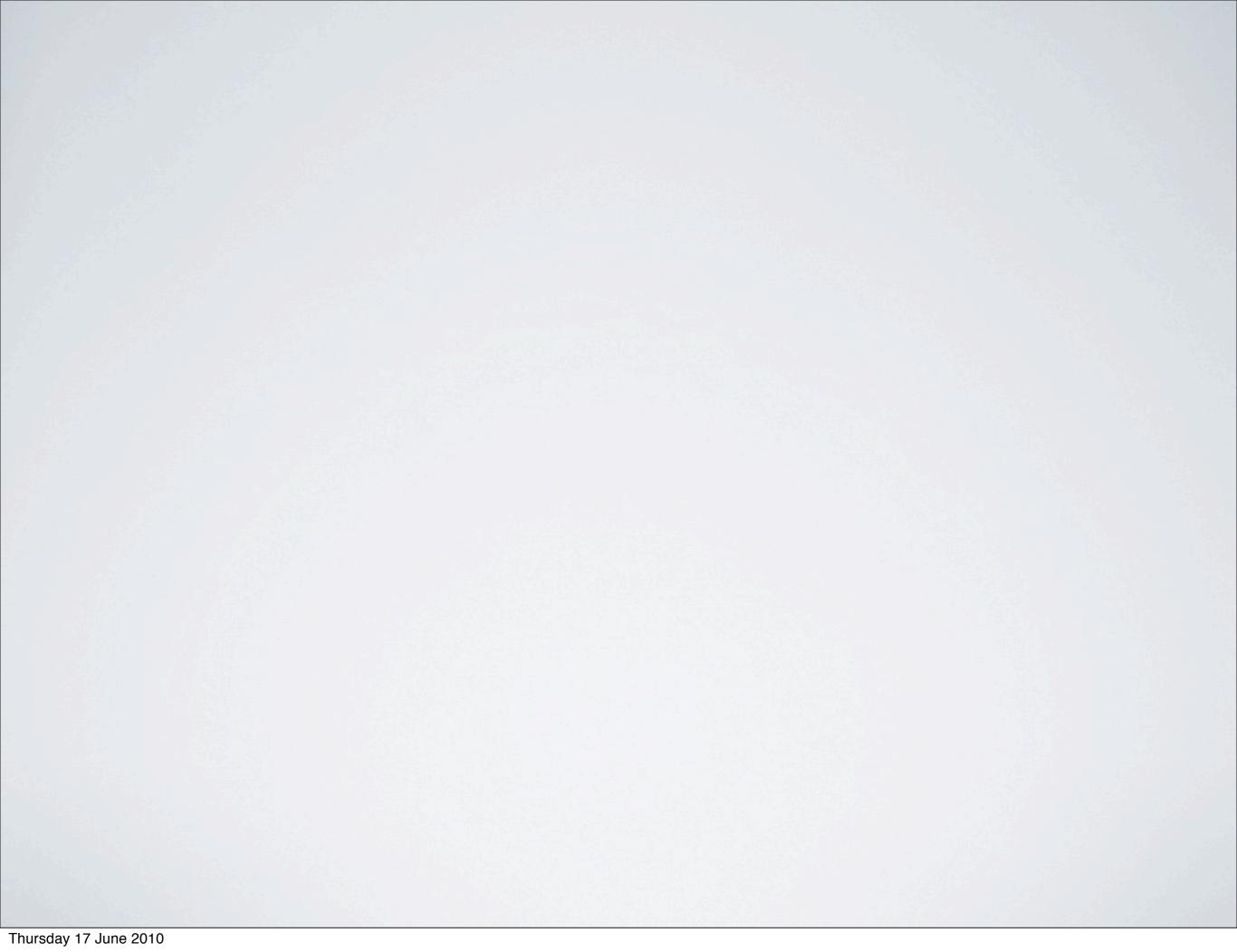


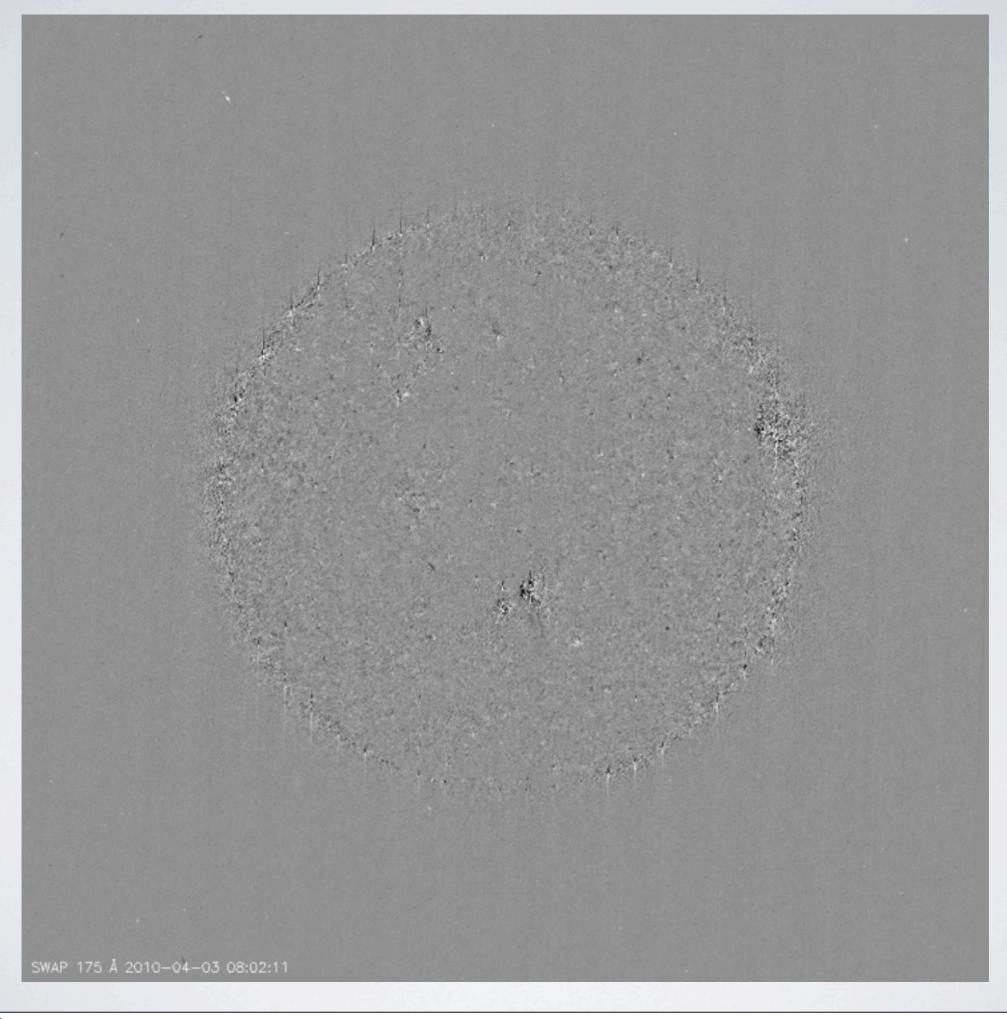
ERUPTION & FLARE

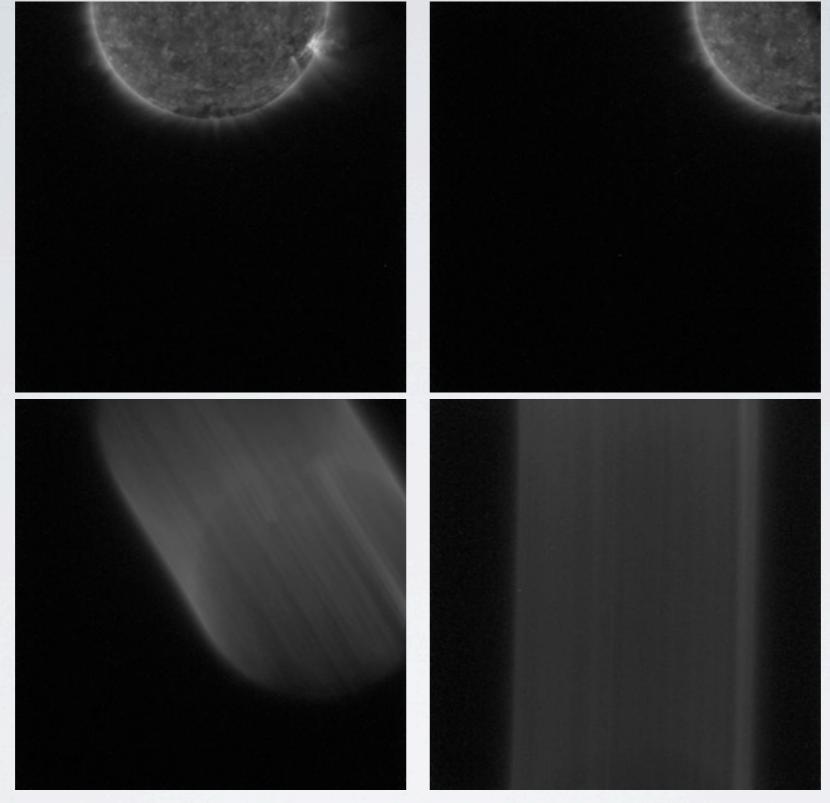
3 April 2010, 09:30 UTC * B7.4 Flare * Geoeffective CME





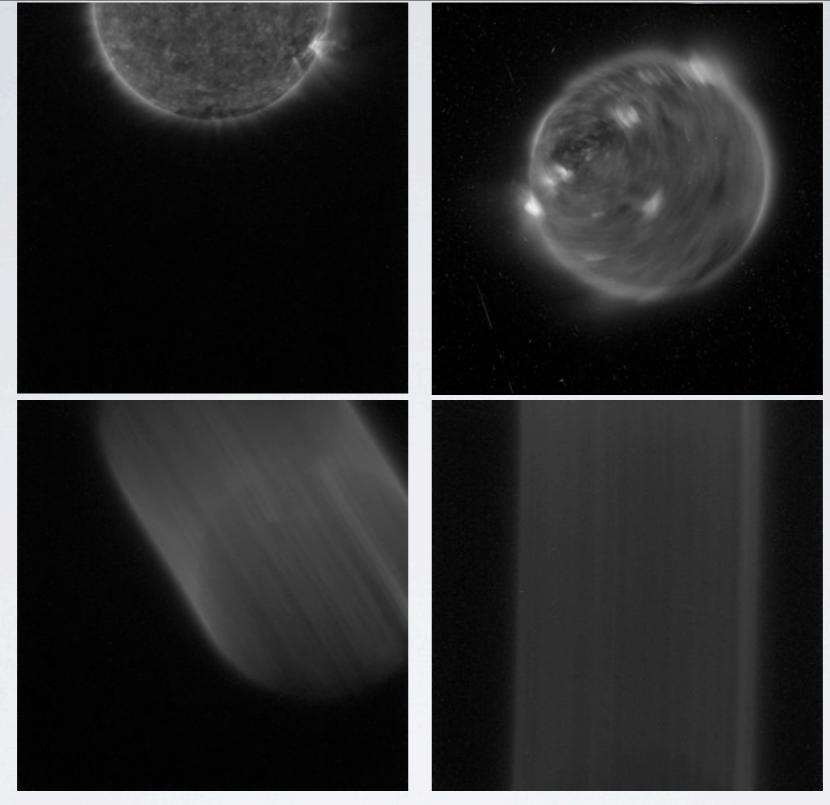






OFF-POINTING

up to now mostly interesting for calibration could be used for CME tracking



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OPEN DATA POLICY

Data are freely available to all users on http://proba2.sidc.be/swap/data http://proba2.sidc.be/lyra/data

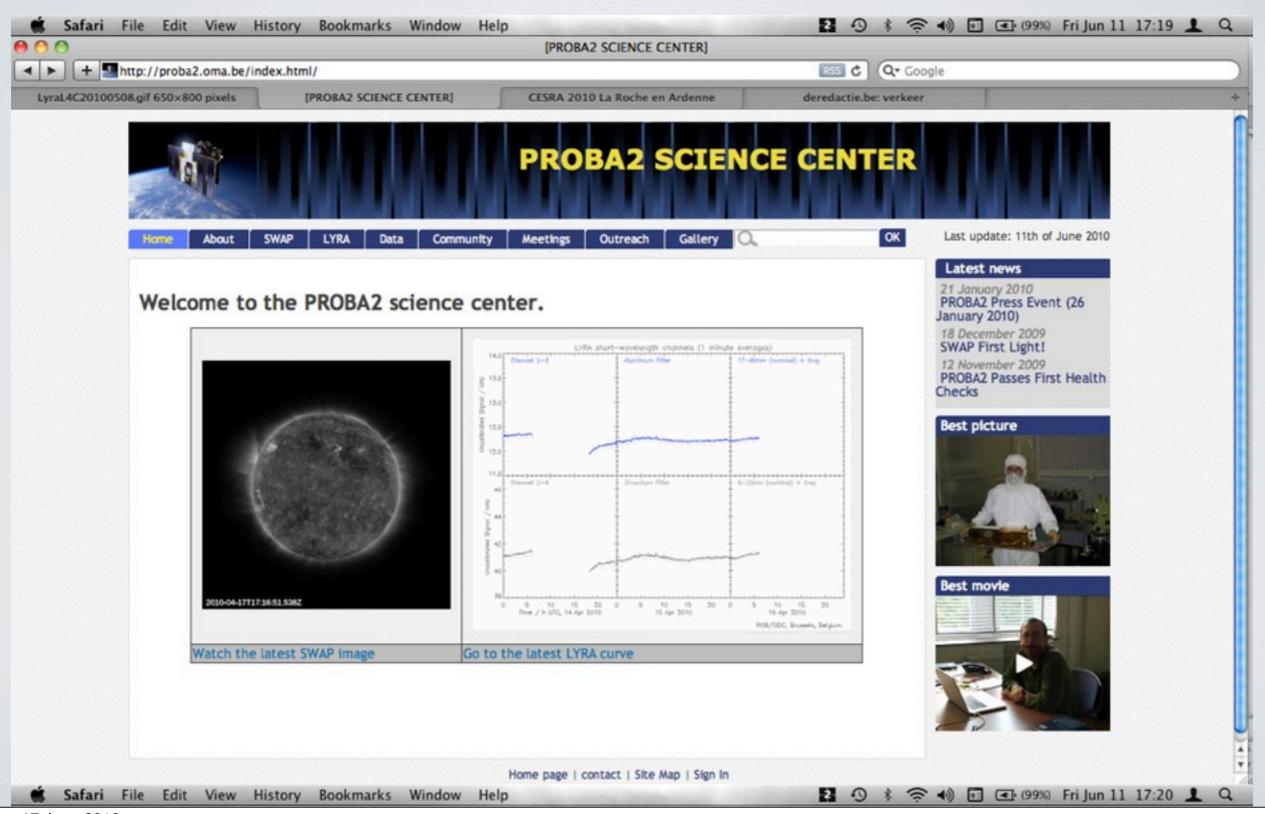
All data ordered in year/month/day folders Fancy data browser to come

Raw Engineering FITS: reformatted, decompressed, long header Base Science Data FITS: (preliminary) calibrated, science header PNG & EPS files: for quicklook purposes SWAP daily movies: for quicklook purposes

SSW will have software trees SWAP & LYRA very soon

FOR MORE INFORMATION

http://proba2.sidc.be/



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