P2SC-ROB-WR-722 - 20240122	P2SC Weekly report	**** ****
Period covered: Date:	′	Royal Observatory of Belgium
Written by: Approved by:		PROBA2 Science Center
To:	LYRA PI, marie.dominique@sidc.be SWAP PI, elke.dhuys@sidc.be	https://proba2.sidc.be ++ 32 (0) 2 3730559
CC:	ROB DIR, ronald@oma.be ESA Redu, Rene.Wittmann@esa.int and Marcus.De.Deus.Silva@esa.int ESA D/SRE, Joe.Zender@esa.int ESA D/TEC, Juha-Pekka.Luntama@esa.int and Melanie.Heil@esa.int	

1. Science

Solar & Space weather events

The level of solar activity¹ fluctuated between **low and moderate** this week.

Only M- and X-flares are mentioned, the most energetic one(s) per day are presented in **bold**:

	Monday 22 Jan	Tuesday 23 Jan	Wednesday 24 Jan	Thursday 25 Jan	Friday 26 Jan	Saturday 27 Jan	Sunday 28 Jan
Activity	moderate	moderate	moderate	low	low	low	low
Flares	M2.1, M1.6, M3.4, M2.0, M1.2, M1.5	M1.0, M1.1, M4.3, M1.3, M1.0, M1.0, M2.3, M2.4, M5.1	M1.3 M1.4 M2.6 M1.0	-	-	-	-

¹ See appendix. All timings are given in UT.

Solar Activity

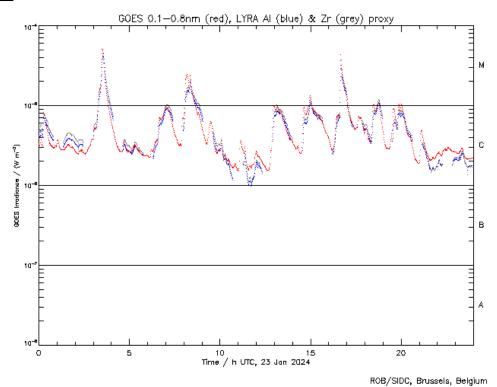
Solar flare activity fluctuated from low to moderate during the week. In order to view the activity of this week in more detail, we suggest to go to the following website from which all the daily (normal and difference) movies can be accessed: https://proba2.oma.be/ssa
This page also lists the recorded flaring events.

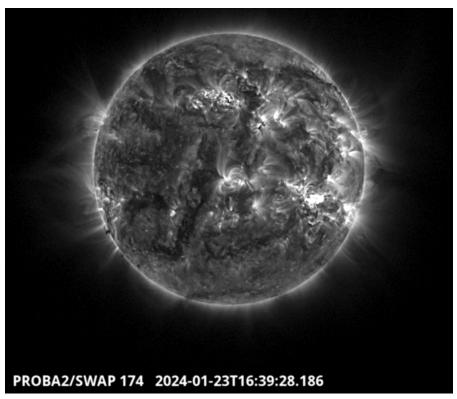
A weekly overview movie can be found here (SWAP week 722).

Details about some of this week's events can be found further below.

If any of the linked movies are unavailable they can be found in the P2SC movie repository here

Tuesday Jan 23





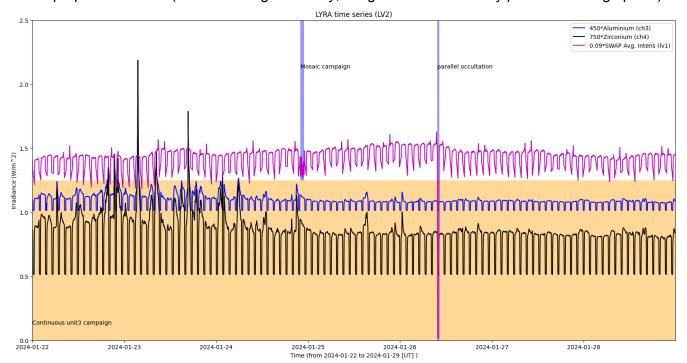
The largest flare of this week was an M5.1, but after a careful inspection of the SWAP data, it is likely that this flux resulted from a summed contribution of two nearly simultaneous flaring regions, NOAA 3559 and 3561. Therefore, the second largest flare was an M4.3, and it was observed by LYRA (top panel) and SWAP (bottom panel). The flare peaked on 2024-Jan-23 at 16:40 UT. It occurred in the south-western quadrant, and it originated from NOAA AR3561.

Find a SWAP movie of the event here.

An overview of the weekly LYRA & SWAP data is provided below:

The following curves are visible:

- black: Zirconium Channel LYRA Unit 2
- blue: Aluminium Channel of LYRA Unit 2
- purple: SWAVINT (SWAP Average Intensity; integrated solar intensity per SWAP image pixel)



Operations and Calibrations:

The blue shaded periods related to SWAP, correspond to, from left to right:

- Mosaic campaign, 2024-Jan-24
- SWAP and LYRA parallel occultation, 2024-Jan-26

The orange shaded periods related to LYRA correspond to, from left to right:

• Unit 3 continuous campaign

The red shaded periods related to other issues corresponds to:

None

2. LYRA instrument status

IOS

Start IOS	Mon Jan 22 2024	LYIOS01051
End IOS	Sun Jan 28 2024	LYIOS01051

LYRA detector temperature

LYRA detector 2 temperature globally varied between 47.3 and 51.1 $^{\circ}\text{C}.$

3. SWAP instrument status

MCPM errors

The number of MCPM recoverable errors increased from 51763 to 51856.

The number of MCPM unrecoverable errors remained at 3135.

IOS

Start IOS	Mon Jan 22 2024	IOS01174
End IOS	Sun Jan 28 2024	IOS01175

SWAP detector temperature

The SWAP Cold Finger Temperature globally varied between -1.77 and 0.87 °C.

4. PROBA2 Science Center Status

The following changes were made to the P2SC:

• None.

5. Data reception & discussions with MOC

Passes

The delivery of the passes for this week (passes 46211 to 46269) was nominal, except for:

• 46214

Data coverage HK

All HK data files (LYRA_AD) have been received, except:

- None.
- LYRA_AD for pass 46214 (from 22 Jan, re-extracted on 23 and 24 Jan) contained many empty files from previous passes. The pipeline could not process the empty files, but the data from pass 46214 had been correctly processed. So no data gap occurred.

Data coverage SWAP

All SWAP Science data files (BINSWAP) have been received, except:

None

Total number of images between 2024 Jan 22 00:00 UT and 2024 Jan 29 00:00 UT: 4179

Highest cadence in this period: 18 seconds

Average cadence in this period: 144.75 seconds Number of image gaps larger than 300 seconds: 176

Largest data gap: 29.48 minutes

Data coverage LYRA

All LYRA Science data files (BINLYRA) have been received, except:

None

6. APPENDIX: Frequently used acronyms

ADPMS Advanced Data and Power Management System

AOCS Attitude and Orbit Control System

APS Active Pixel image Sensor

ASIC Application Specific Integrated Circuit

BBE Base Band Equipment
CME Coronal Mass Ejection

COGEX Cool Gas Generator Experiment

CRC Cyclic Redundancy Check
DAC Data Acquisition Controller
DBR Deployment, backup & recovery
DDA Decommutated data archive
ESP Experimental Solar Panel

FITS Flexible Image Transport System

FOV Field Of View FPA Focal Plane Assembly

FPGA Field Programmable Gate Arrays

GPS Global Positioning System

HK Housekeeping

IOS Instrument Operations Sheet

LED Light Emitting Diode
LYRA LYman alpha RAdiometer

LYTMR LYRA Telemetry Reformatter (software module of P2SC)
LYEDG LYRA Engineering Data Generator (software module of P2SC)

MCPM Mass Memory, Compression and Packetisation Module

MOC Mission Operation Center NDR Non Destructive Readout

OBSW On board Software
PI Principal Investigator
P2SC PROBA2 Science Center
ROB Royal Observatory of Belgium

SAA South Atlantic Anomaly
SEU Single Event Upset

SoFAST | Solar Feature Automated Search Tool

SWAP Sun Watcher using APS detector and image Processing

SWAVINT | SWAP AVerage INTensity

SWBSDG | SWAP Base Science Data Generator

SWEDG SWAP Engineering Data Generator (software module of P2SC)
SWTMR SWAP Telemetry Reformatter (software module of P2SC)

TBC To Be Confirmed

TBD To Be Defined TC Telecommand

UTC Coordinated Universal Time

UV Ultraviolet

VFC Voltage to Frequency Converter

7. APPENDIX Solar Activity Definitions

In the science section we use the following solar activity standards.

The standard scale for solar activity is:

- very low (almost no flares, only B)
- low (a few C flares)
- moderate (many C flares and at least an M flare)
- high (several M flares and an X flare)
- very high (continuous background of C flares, numerous M flares, more than one X flare)