P2SC-ROB-WR-412 - 20180212 Weekly report #412	P2SC Weekly report	****
Period covered: Date: Written by: Approved by:	<b>'</b>	Royal Observatory of Belgium - PROBA2 Science Center
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## 1. Science

## **Solar & Space weather events**

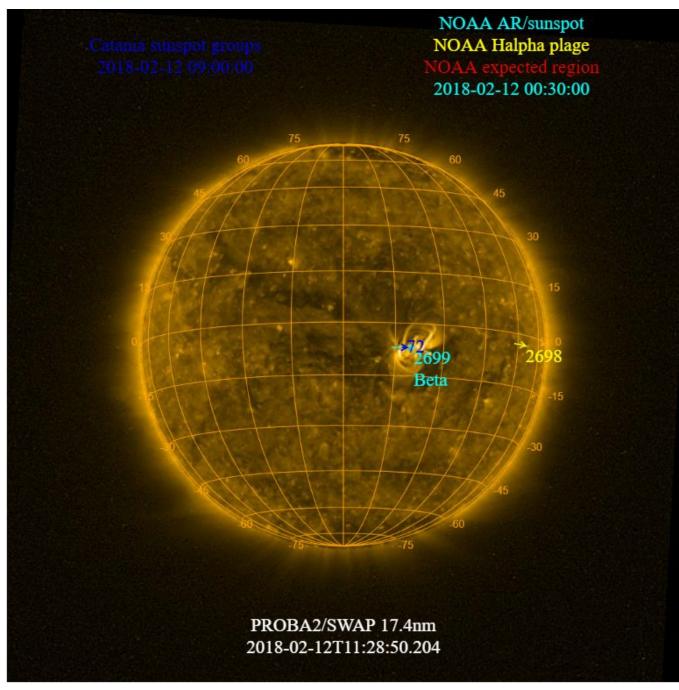
The level of solar activity¹ fluctuated between **very low and low** this week.

Only M- and X-flares are mentioned, the most energetic one(s) per day are presented in **bold**:

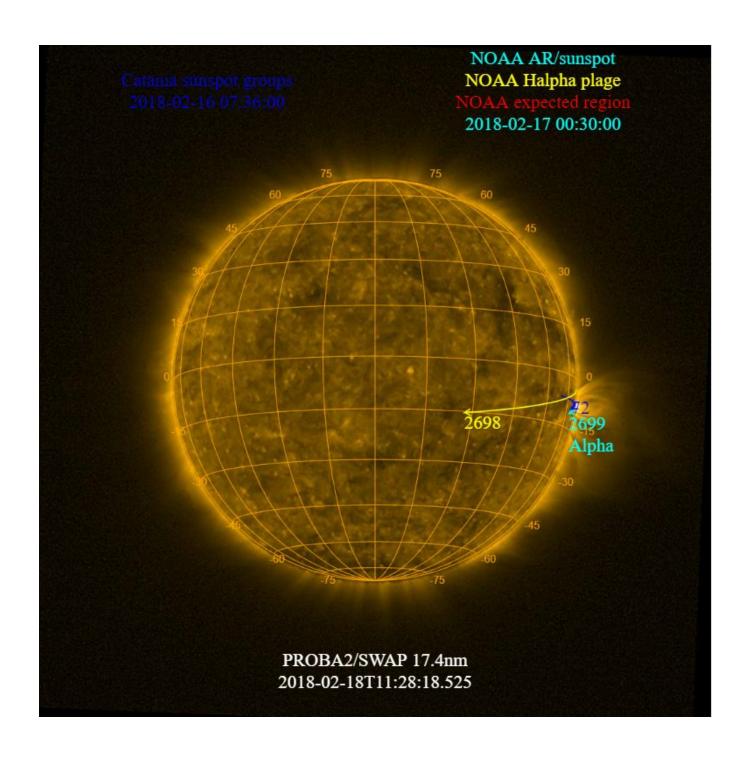
	Monday 12 Feb	Tuesday 13 Feb	Wednesday 14 Feb	Thursday 15 Feb	Friday 16 Feb	Saturday 17 Feb	Sunday 18 Feb
Activity	low	very low	very low	very low	very low	very low	very low
Flares	-	-	-	-	-	-	-

<sup>&</sup>lt;sup>1</sup> See appendix. All timings are given in UT.

The SWAP images of Feb 12 and Feb 18 are shown below, with annotated active regions.



http://sidc.be/soteria/soteria.php



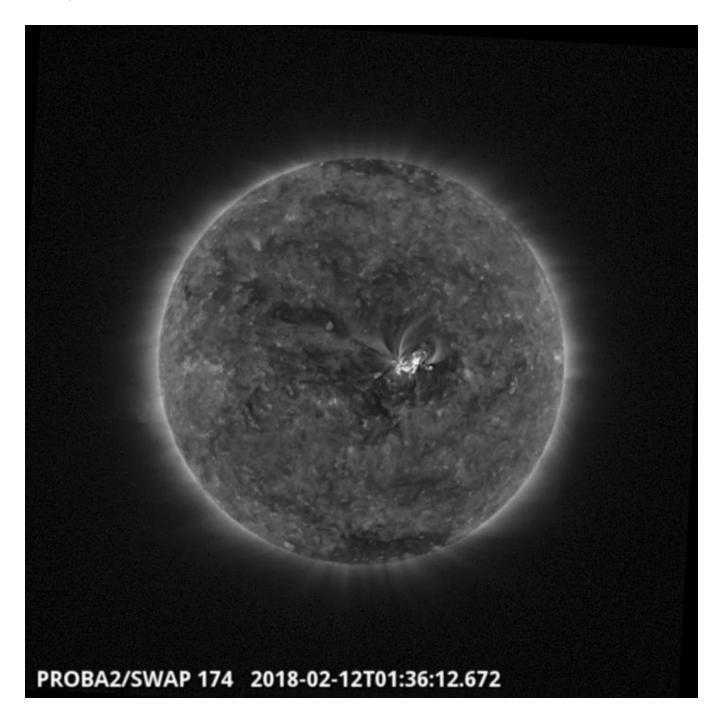
#### **Solar Activity**

Solar flare activity fluctuated between very low and low during the week. In order to view the activity of this week in more detail, we suggest to go to the following website from which all the daily (normal and difference) movies can be accessed: <a href="http://proba2.oma.be/ssa">http://proba2.oma.be/ssa</a>
This page also lists the recorded flaring events.

A weekly overview movie can be found here (SWAP week 412).

Details about some of this week's events, can be found further below.

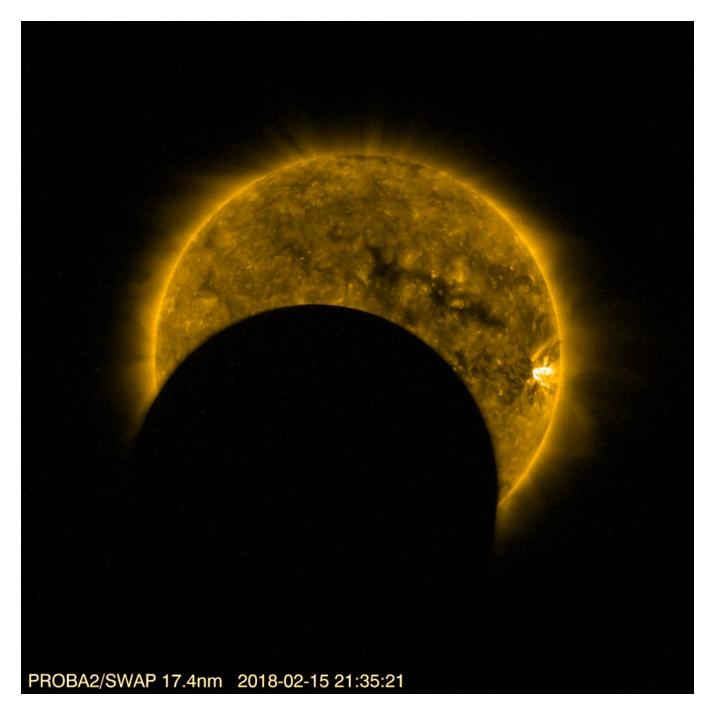
If any of the linked movies are unavailable they can be found in the P2SC movie repository here



A C-class (C1.5) flare, associated with NOAA AR 2699, was observed by SWAP on 2018-Feb-12. The flare is visible near the solar disk center in the SWAP image above at 01:36 UT. This flare was followed by several B-class flares throughout the week.

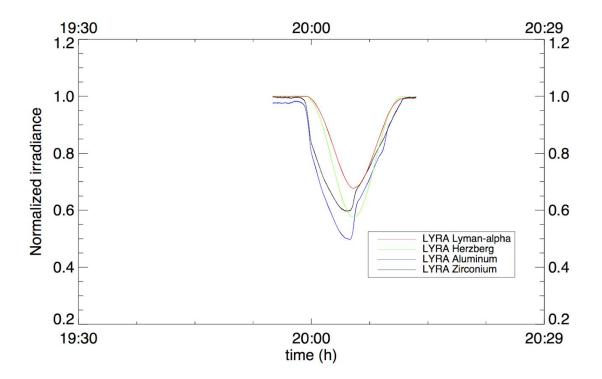
Find a movie of the events **here** (SWAP movie)

## Thursday Feb 15



On February 15, two eclipses occured (between 19:59 and 20:11 UT, and between 21:28 and 21:41 UT respectively). The SWAP image above was taken during the second one. Additionally, the Moon appeared two more times in the SWAP and LYRA field of view, but without obscuring the solar disk.

Find a movie of the events <a href="https://example.com/here">here</a> (SWAP movie)

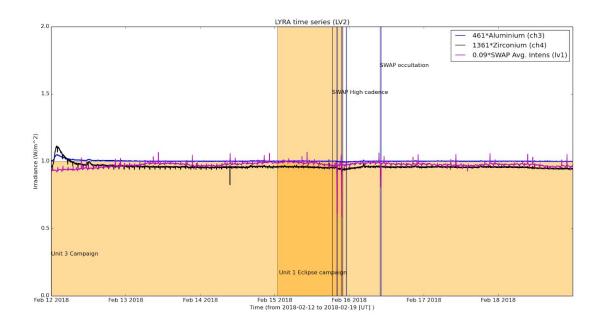


On February 15, the first partial eclipse (between 19:59 and 20:11 UT) was observed by LYRA, the above image shows the irradiance curves from LYRA throughout the event. Find LYRA fits files for this event <a href="https://example.com/here">here</a>.

An overview of the weekly LYRA & SWAP data is provided below:

The following curves are visible:

- black: Zirconium Channel LYRA Unit 2
- blue: Aluminium Channel of LYRA Unit 2
- purple: SWAVINT (SWAP Average Intensity; integrated solar intensity per SWAP image pixel)



The blue shaded periods related to SWAP, correspond to, from left to right:

- Eclipse high cadence campaign, 2018-Feb-15 (between 18:28 UT and 18:38 UT, between 19:56 and 20:16 UT, between 21:23 UT and 21:44 UT and between 23:00 UT and 23:12 UT respectively)
- Parallel occultation campaign with LYRA, 2018-Feb-16

The orange shaded periods related to LYRA correspond to, from left to right:

- Continuous Unit 3 campaign, from 2018-Feb-12 to 2018-Feb-15 00:40 and from 2018-Feb-15 22:15 to 2018-Feb-18
- LYRA (Unit1) eclipse campaign, 2018-Feb-15 from 00:53 UT to 22:00 UT

The red shaded periods related to other issues corresponds to:

None

#### Outreach, papers, presentations, etc.

Please consult <a href="http://proba2.oma.be/science/publications">http://proba2.oma.be/science/publications</a> for a list of interesting articles using SWAP & LYRA data, as well as a link to the complete article list.

The science section of this weekly report is also published in the weekly STCE newsletter (http://www.stce.be/newsletter/newsletter.php).

"The detection of ultra-relativistic electrons in low Earth orbit" by Katsiyannis et al. In this paper the authors looked at relativistic electrons with an energy range of 2 to 8 MeV present at high latitudes in the Earth atmosphere. The particles were indirectly detected by LYRA as a background signal. Combining LYRA observations with energetic particle telescope (EPT) observations the authors showed that the observed events are strongly correlated to geomagnetic activity and appear even during modest disturbances.

"On the Performance of Multi-Instrument Solar Flare Observations During Solar Cycle 24" by Milligan and Ireland. In this paper the authors describe a technique that combines observations from various solar missions instruments. As part of the dataset, SWAP's large field of view was used to produce the context image to show the location of a flare detected by all instruments and to emphasize the importance of large FOVs in such observations.

"Vertical Thermospheric Density Profiles From EUV Solar Occultations Made by PROBA2 LYRA for Solar Cycle 24" by Thiemann et al. In this paper the authors develop a new data set of summed neutral nitrogen and oxygen number density profiles of the Earths atmosphere, spanning altitudes between 150 and 400 km. The neutral density profiles are derived from solar occultation measurements made by the Zirconium channel on LYRA. The observations are compared and match well with model made predictions. The density profiles are used to characterize the response of thermospheric density to solar EUV irradiance variability. LYRA is used through out the paper, making use of occultation observations. When PROBA2 passes behind the Earth in its orbit, LYRA is used to look at how the EUV emission of the Sun is extinguished. Note this paper was published in 2017, but was only listed on NASA ADS in the last weeks.

"Modelling Quasi-Periodic Pulsations in Solar and Stellar Flares" by McLaughlin et al was published in Space Science Reviews. This paper focuses on the radiation generated in solar and stellar flares, which show a pronounced oscillatory pattern, with characteristic periods ranging from a fraction of a second to several minutes. LYRAs high cadence observations are at the forefront of QPP observations due to LYRAs high cadence observations, and are discussed throughout this paper. An example is given in Figure 1 for the X4.9 flare of 25 February 2014. QPP oscillations with a period of ~35 s are clearly visible as an oscillatory train in all displayed time series.

#### **Guest Investigator Program**

Ed Thiemann from LASP, Boulder, Colorado is visiting the P2SC for the week between Feb-08 and Feb-15 to begin his collaboration on the project "Comparing the response of the thermospheres of Earth and Mars to solar forcing with contemporaneous solar EUV occultations." During his visit he also gave a STCE seminar entitled "Mars Thermospheric Variability Revealed by MAVEN EUVM Solar Occultations".

## 2. LYRA instrument status

#### Calibration

No calibration campaign on this week.

## IOS & operations

Monday 12 Feb	Tuesday 13 Feb	Wednesday 14 Feb	Thursday 15 Feb	Friday 16 Feb	Saturday 17 Feb	Sunday 18 Feb
Nominal acquisition + U3	Nominal acquisition + U3	Nominal acquisition + U3	Nominal acquisition + U3	Nominal acquisition + U3+ U1 eclipse campaign	Nominal acquisition + U3	Nominal acquisition + U3
LYIOS00674	LYIOS00674	LYIOS00675	LYIOS00675	LYIOS00676	LYIOS00676	LYIOS00676

The following science campaigns were performed by LYRA:

Continuous Unit 3 campaign

On 2018-Feb-15:

• LYRA (Unit1) eclipse campaign

## LYRA detector temperature

LYRA detector 2 temperature globally varied between 53.33 and 55.35 °C.

## 3. SWAP instrument status

#### Calibration

No Calibration campaign on this week.

#### **MCPM errors**

The number of MCPM recoverable errors increased from 1638 to 1951.

The number of MCPM unrecoverable errors remained at 0.

## **IOS & operations**

Monday	Tuesday	Wednesday	Thursday	Friday	Saturday	Sunday
12 Feb	13 Feb	14 Feb	15 Feb	16 Feb	17 Feb	18 Feb
Nominal acquisition	Nominal acquisition + calibration	Nominal acquisition				
IOS00761	IOS00761	IOS00762	IOS00762	IOS00762	IOS00763	IOS00763
609 images	728 images	586 images	889 images	775 images	680 images	597 images

Special operations for SWAP, this week:

On 2018-Feb-15:

• Eclipse high cadence campaign

On 2018-Feb-16:

Occultation campaign

#### **SWAP** detector temperature

The SWAP Cold Finger Temperature globally varied between 2.07 and 3.91 °C.

## 4. PROBA2 Science Center Status

The main operator is Laurence Wauters.

The following changes were made to the P2SC:

• None.

## 5. Data reception & discussions with MOC

#### **Passes**

The delivery of the passes for this week (passes 26601 to 26668) was nominal, except for:

• None.

#### Data coverage HK

All HK data files (LYRA\_AD) have been received, except:

None.

### **Data coverage SWAP**

All SWAP Science data files (BINSWAP) have been received, except:

None.

Total number of images between 2018 Feb 12 00:00 UT and 2018 Feb 19 00:00 UT: 4925

Highest cadence in this period: 0 seconds

Average cadence in this period: 122.79 seconds Number of image gaps larger than 300 seconds: 126

Largest data gap: 18.33 minutes

#### **Data coverage LYRA**

All LYRA Science data files (BINLYRA) have been received, except:

None

## 6. APPENDIX: Frequently used acronyms

ADPMS Advanced Data and Power Management System

AOCS Attitude and Orbit Control System

APS Active Pixel image Sensor

ASIC Application Specific Integrated Circuit

BBE Base Band Equipment
CME Coronal Mass Ejection

COGEX Cool Gas Generator Experiment

CRC Cyclic Redundancy Check
DAC Data Acquisition Controller
DBR Deployment, backup & recovery
DDA Decommutated data archive

ESP Experimental Solar Panel

FITS Flexible Image Transport System

FOV Field Of View FPA Focal Plane Assembly

FPGA Field Programmable Gate Arrays

GPS Global Positioning System

HK Housekeeping

IOS Instrument Operations Sheet

LED Light Emitting Diode

LYRA LYman alpha RAdiometer

LYTMR LYRA Telemetry Reformatter (software module of P2SC)
LYEDG LYRA Engineering Data Generator (software module of P2SC)

MCPM Mass Memory, Compression and Packetisation Module

MOC Mission Operation Center NDR Non Destructive Readout

OBSW On board Software
PI Principal Investigator
P2SC PROBA2 Science Center
ROB Royal Observatory of Belgium

SAA South Atlantic Anomaly
SEU Single Event Upset

SoFAST | Solar Feature Automated Search Tool

SWAP Sun Watcher using APS detector and image Processing

SWAVINT | SWAP AVerage INTensity

SWBSDG SWAP Base Science Data Generator

SWEDG SWAP Engineering Data Generator (software module of P2SC)
SWTMR SWAP Telemetry Reformatter (software module of P2SC)

TBC To Be Confirmed
TBD To Be Defined
TC Telecommand

UTC Coordinated Universal Time

UV Ultraviolet

VFC Voltage to Frequency Converter

# 7. APPENDIX Solar Activity Definitions

In the science section we use the following solar activity standards.

The standard scale for solar activity is:

- very low (almost no flares, only B)
- low (a few C flares)
- moderate (many C flares and at least an M flare)
- high (several M flares and an X flare)
- very high (continuous background of C flares, numerous M flares, more than one X flare)