P2SC-ROB-WR-359 - 20170206 Weekly report #359	P2SC Weekly report	* **** ****
Period covered: Date:	Mon Feb 06 to Sun Feb 12, 2017 14 Feb 2017	Royal Observatory of Belgium
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1. Science

Solar & Space weather events

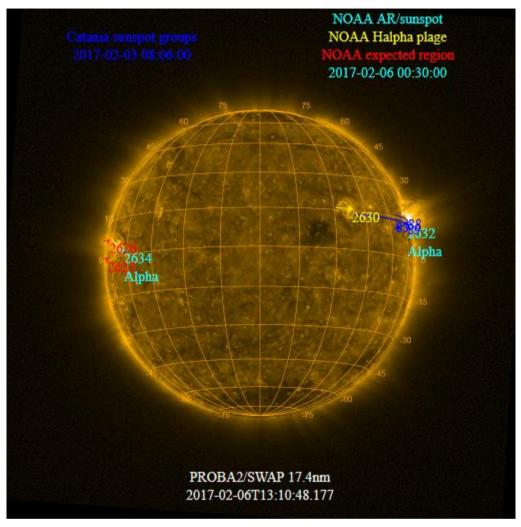
The level of solar activity¹ fluctuated between **very low and low** this week.

Only M- and X-flares are mentioned, the most energetic one(s) per day are presented in **bold**:

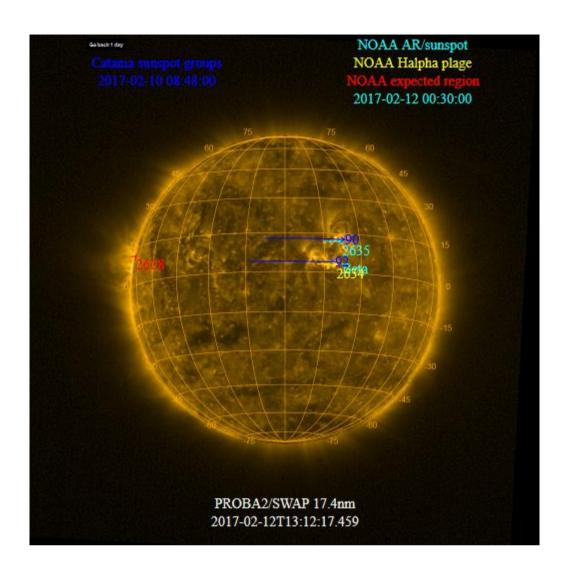
	Monday 06 Feb	Tuesday 07 Feb	Wednesday 08 Feb	Thursday 09 Feb	Friday 10 Feb	Saturday 11 Feb	Sunday 12 Feb
Activity	very low	very low	very low	low	very low	very low	very low
Flares	-	-	-	-	-	-	-

¹ See appendix. All timings are given in UT.

The SWAP images of Feb 06 and Feb 12 are shown below, with annotated active regions.



http://sidc.be/soteria/soteria.php



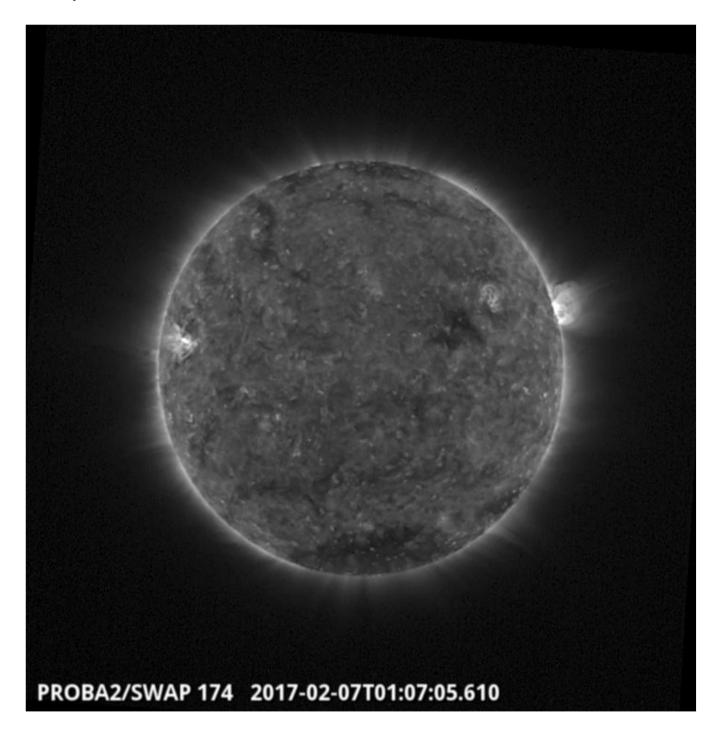
Solar Activity

Solar flare activity fluctuated between very low and low during the week. In order to view the activity of this week in more detail, we suggest to go to the following website from which all the daily (normal and difference) movies can be accessed: http://proba2.oma.be/ssa
This page also lists the recorded flaring events.

A weekly overview movie can be found here (SWAP week 359).

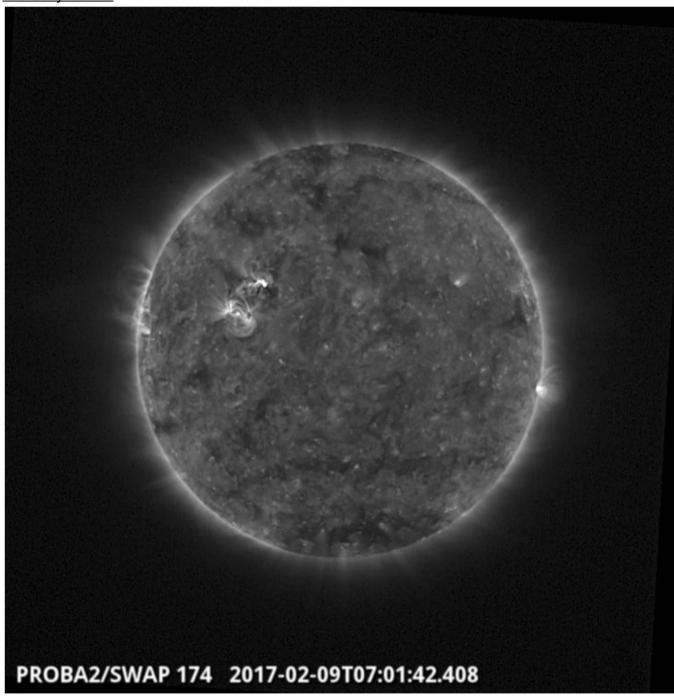
Details about some of this week's events, can be found further below.

If any of the linked movies are unavailable they can be found in the P2SC movie repository here



A Southern coronal hole, which produced enhanced geomagnetic conditions at the Earth was seen on the the solar disk at the beginning of the week, this can be seen in the above SWAP image in the Western hemisphere.

Find a movie of the events here (SWAP movie)



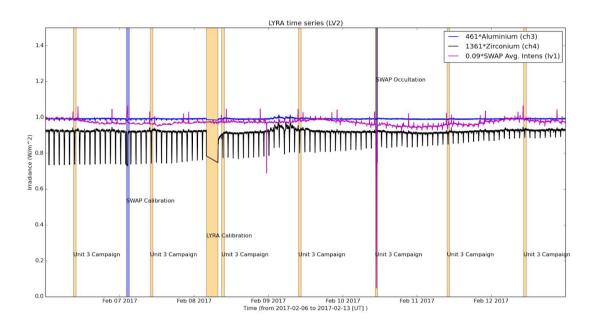
AR 2635 produced eleven B-flares and three C-class flares on 2017-Feb-09 (the largest one was a C1.3). This active region is visible above in the North-East part of the Sun flaring at 07:02 UT when it produced a C1.1 class flare.

Find a movie of the events here (SWAP movie)

An overview of the weekly LYRA & SWAP data is provided below:

The following curves are visible:

- black: Zirconium Channel LYRA Unit 2
- blue: Aluminium Channel of LYRA Unit 2
- purple: SWAVINT (SWAP Average Intensity; integrated solar intensity per SWAP image pixel)



The blue shaded periods correspond to, from left to right:

- SWAP calibration, Feb 07
- SWAP occultation campaign, Feb 10

The orange shaded periods correspond to, from left to right:

- Daily unit 3 campaign, 2017-Feb 06
- Daily unit 3 campaign, 2017-Feb 07
- LYRA bi-weekly Calibration, 2017-Feb-08
- Daily unit 3 campaign, 2017-Feb-08
- Daily unit 3 campaign, 2017-Feb-09
- Daily unit 3 campaign, 2017-Feb-10
- Daily unit 3 campaign, 2017-Feb-11
- Daily unit 3 campaign, 2017-Feb-12

The red shaded period corresponds to:

None

Outreach, papers, presentations, etc.

Please consult http://proba2.oma.be/science/publications for a list of interesting articles using SWAP & LYRA data, as well as a link to the complete article list.

The science section of this weekly report is also published in the weekly STCE newsletter (http://www.stce.be/newsletter/newsletter.php).

Schanche et al. 2016, published a paper entitled: "The blob connection: Searching for low coronal signatures of solar post-CME blobs". Following coronal mass ejections or other eruptions on the Sun, it's postulated that the magnetic field trailing the eruption collapses in on itself due to magnetic pressures, this can result in a magnetic null point trailing the CME. Magnetic null points, although invisible, can be tracked through the ejection of material from the null point. This material has previously been interpreted as blobs of material that are seen to travel from the null point, along the current sheet trailing the eruption, into the lower corona. The authors use PROBA2 SWAP and SDO AIA observations to try and track these blobs in the lower corona, but fail in both EUV imagers. This null result has led the authors to question the previously estimated properties of the blobs, suggesting that they form higher in the solar atmosphere, or do not coalesce as we would expect and / or have lower densities than previously anticipated.

Guest Investigator Program

None

2. LYRA instrument status

Calibration

Calibration campaign on Wednesday this week.

IOS & operations

Monday 06 Feb	Tuesday 07 Feb	Wednesday 08 Feb	Thursday 09 Feb	Friday 10 Feb	Saturday 11 Feb	Sunday 12 Feb
Nominal acquisition + daily U3	Nominal acquisition + daily U3	Nominal acquisition + daily U3+ Calibration	Nominal acquisition + daily U3	Nominal acquisition + daily U3	Nominal acquisition + daily U3	Nominal acquisition + daily U3
LYIOS00599	LYIOS00599	LYIOS00599	LYIOS00599	LYIOS00600	LYIOS00600	LYIOS00600

The following science campaigns were performed by LYRA:

• Daily U3 observations campaign

On 2017-Feb-08

• LYRA bi-weekly calibration, 2017-Feb-08

LYRA detector temperature

LYRA detector 2 temperature globally varied between 51.67 and 54.48 °C.

3. SWAP instrument status

Calibration

Calibration campaign on Tuesday this week.

MCPM errors

The number of MCPM recoverable errors increased from 6021 and 6254.

The number of MCPM unrecoverable errors remained at 0.

IOS & operations

Monday	Tuesday	Wednesday	Thursday	Friday	Saturday	Sunday
06 Feb	07 Feb	08 Feb	09 Feb	10 Feb	11 Feb	12 Feb
Nominal acquisition	Nominal acquisition+ Calibration	Nominal acquisition	Nominal acquisition	Nominal acquisition+ Occultation	Nominal acquisition	Nominal acquisition
IOS00685	IOS00686	IOS00686	IOS00686	IOS00687	IOS00687	IOS00687
749 images	691 images	780 images	717 images	746 images	671 images	650 images

Special operations for SWAP, this week:

- SWAP calibration, 2017-Feb-07
- SWAP occultation campaign, 2017-Feb-10

SWAP detector temperature

The SWAP Cold Finger Temperature globally varied between 2.23 and 4.31 °C.

4. PROBA2 Science Center Status

The main operator is Laurence Wauters.

The following changes were made to the P2SC:

• None.

5. Data reception & discussions with MOC

Passes

The delivery of the passes for this week (passes 23118 to 23184) was nominal, except for:

None.

Data coverage HK

All HK data files (LYRA_AD) have been received, except:

None.

Data coverage SWAP

All SWAP Science data files (BINSWAP) have been received, except:

None.

Total number of images between 2017 Feb 06 0UT and 2017 Feb 13 0UT: 5004

Highest cadence in this period: 29 seconds

Average cadence in this period: 120.84 seconds Number of image gaps larger than 300 seconds: 122

Largest data gap: 27.68 minutes

Data coverage LYRA

All LYRA Science data files (BINLYRA) have been received, except:

- None
- On Feb 12, two files of the BINLYRA.tar for pass 23177 were corrupted.

6. APPENDIX: Frequently used acronyms

ADPMS Advanced Data and Power Management System

AOCS Attitude and Orbit Control System

APS Active Pixel image Sensor

ASIC Application Specific Integrated Circuit

BBE Base Band Equipment CME Coronal Mass Ejection

COGEX Cool Gas Generator Experiment

CRC Cyclic Redundancy Check
DAC Data Acquisition Controller

DBR Deployment, backup & recovery
DDA Decommutated data archive
ESP Experimental Solar Panel

FITS Flexible Image Transport System

FOV Field Of View FPA Focal Plane Assembly

FPGA Field Programmable Gate Arrays

GPS Global Positioning System

HK Housekeeping

IOS Instrument Operations Sheet

LED Light Emitting Diode

LYRA LYman alpha RAdiometer

LYTMR LYRA Telemetry Reformatter (software module of P2SC)
LYEDG LYRA Engineering Data Generator (software module of P2SC)

MCPM Mass Memory, Compression and Packetisation Module

MOC Mission Operation Center NDR Non Destructive Readout

OBSW On board Software
PI Principal Investigator
P2SC PROBA2 Science Center
ROB Royal Observatory of Belgium

SAA South Atlantic Anomaly
SEU Single Event Upset

SoFAST Solar Feature Automated Search Tool

SWAP Sun Watcher using APS detector and image Processing

SWAVINT | SWAP AVerage INTensity

SWBSDG SWAP Base Science Data Generator

SWEDG SWAP Engineering Data Generator (software module of P2SC)
SWTMR SWAP Telemetry Reformatter (software module of P2SC)

TBC To Be Confirmed
TD To Be Defined
TC Telecommand

UTC Coordinated Universal Time

UV Ultraviolet

VFC Voltage to Frequency Converter

7. APPENDIX Solar Activity Definitions

In the science section we use the following solar activity standards.

The standard scale for solar activity is:

- very low (almost no flares, only B)
- low (a few C flares)
- moderate (many C flares and at least an M flare)
- high (several M flares and an X flare)
- very high (continuous background of C flares, numerous M flares, more than one X flare)