P2SC-ROB-WR-291 - 20151019 Weekly report #291	P2SC Weekly report	**** <u>***</u>
Period covered: Date:	Mon Oct 19 to Sun Oct 25, 2015 28 Oct 2015	Royal Observatory of Belgium -
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## 1. Science

## Solar & Space weather events

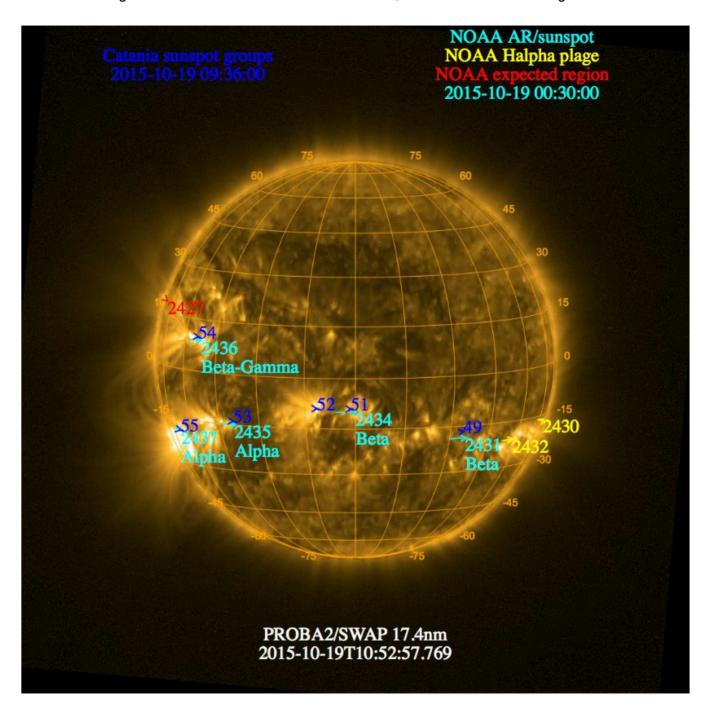
The level of solar activity¹ fluctuated between **very low** and **low** this week.

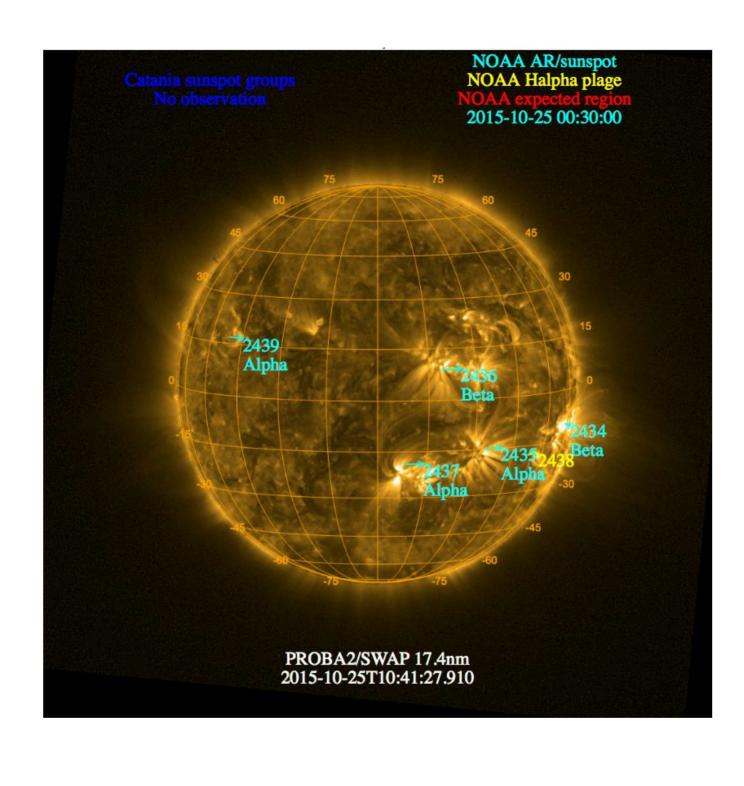
Only M- and X-flares are mentioned, the most energetic one(s) per day are presented in **bold**:

	Monday 19 Oct	Tuesday 20 Oct	Wednesday 21 Oct	Thursday 22 Oct	Friday 23 Oct	Saturday 24 Oct	Sunday 25 Oct
Activity	low	low	low	low	very low	low	very low
Flares	-	-	-	-	-	-	-

<sup>&</sup>lt;sup>1</sup> See appendix. All timings are given in UT.

The SWAP images of Oct 19 and Oct 25 are shown below, with annotated active regions.





### **Solar Activity**

Solar flare activity fluctuated between very low and low during the week.

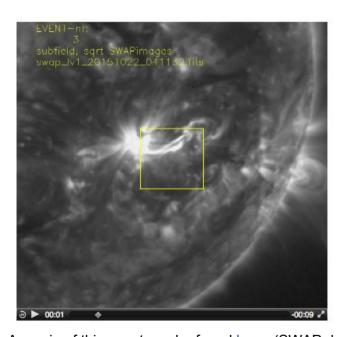
In order to view the activity of this week in more detail, we suggest to go to the following website from which all the daily (normal and difference) SWAP movies can be accessed: <a href="http://proba2.oma.be/ssa">http://proba2.oma.be/ssa</a> This page also lists recorded flaring events.

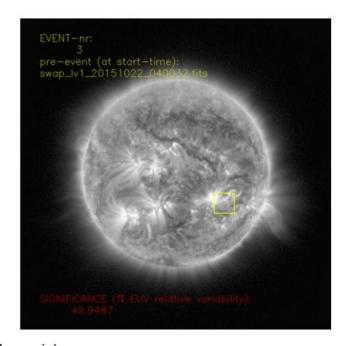
A weekly overview movie can be found <a href="here">here</a> (SWAP week 291).

Solar activity was low this week. AR 2434 produced a long duration C-class flare on 2015-Oct-22 between 03:13 and 05:15 UT. Below we provide SWAP images from the time when this C-class flare occurred. The annotated snapshots are produced by the Solar Feature Automated Search Tool (SoFAST). This tool detects dynamic solar events in EUV images from SWAP in near real-time. The snapshots illustrate the location of the flare on the solar disk (right) and a zoomed image (left).

The complete SoFAST online event list and additional plots are available on <a href="http://www.sidc.be/sofast">http://www.sidc.be/sofast</a>.

2015-Oct-22: Long duration C4.4 flare



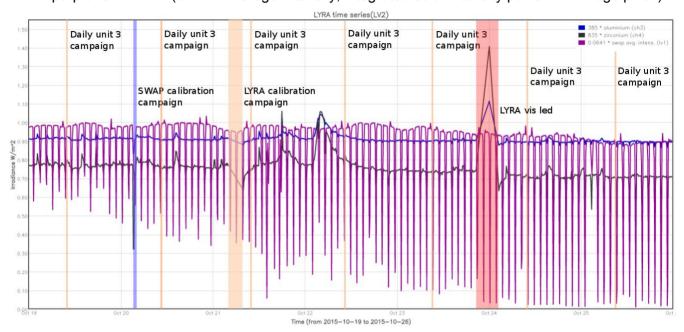


A movie of this event can be found <a href="here">here</a> (SWAP daily movie)

An overview of the weekly LYRA & SWAP data is provided below.

The following curves are visible:

- black: Zirconium Channel LYRA Unit 2
- blue: Aluminium Channel of LYRA Unit 2
- purple: SWAVINT (SWAP Average Intensity; integrated solar intensity per SWAP image pixel)



The blue shaded periods correspond to, from left to right:

SWAP bi-weekly calibration campaign on 2015-Oct-20

The orange shaded periods correspond to, from left to right:

- LYRA daily U3 campaign on 2015-Oct-19
- LYRA daily U3 campaign on 2015-Oct-20
- LYRA short bi-weekly calibration on 2015-Oct-21
- LYRA daily U3 campaign on 2015-Oct-21
- LYRA daily U3 campaign on 2015-Oct-22
- LYRA daily U3 campaign on 2015-Oct-23
- LYRA daily U3 campaign on 2015-Oct-24
- LYRA daily U3 campaign on 2015-Oct-25

The red shaded period corresponds to:

LYRA visual LED switched on between 2015-Oct-23 20:48 and 2015-Oct-24 02:22 UT

### Outreach, papers, presentations, etc.

Please consult <a href="http://proba2.oma.be/science/publications">http://proba2.oma.be/science/publications</a> for a list of interesting articles using SWAP & LYRA data, as well as a link to the complete article list.

The science section of this weekly report is also published in the weekly STCE newsletter (<a href="http://www.stce.be/newsletter/newsletter.php">http://www.stce.be/newsletter/newsletter.php</a>).

 PROBA2 Guest Investigator Michael Kirk gave a presentation of his work on: "Mapping Solar Cycles Through Polar Coronal Holes" at ROB on 2015-Oct-22

### **Guest Investigator Program**

- 2015-Oct-14 until 2015-Oct-23: Michael Kirk is working at the P2SC, using SWAP data to perform research on: "Mapping Solar Cycles Through Polar Coronal Holes."
- 2015-Oct-21 until 2015-Nov-01: Sabrina Savage is working at the P2SC, using SWAP data to tie coronal flows to flare arcades spines.
- 2015-Oct-21 until 2015-Oct-29: Adam Kobelski is working at the P2SC, using SWAP data to study AR-AR reconnection after flux emergence.

### 2. LYRA instrument status

### Calibration

Calibration campaign on Wednesday this week.

### **IOS & operations**

Monday 19 Oct	Tuesday 20 Oct	Wednesday 21 Oct	Thursday 22 Oct	Friday 23 Oct	Saturday 24 Oct	Sunday 25 Oct
Nominal acquisition + daily U3	Nominal acquisition + daily U3	Nominal acquisition + daily U3 + calibration	Nominal acquisition + daily U3	Nominal acquisition + daily U3 + vis led	Nominal acquisition + daily U3	Nominal acquisition + daily U3
LYIOS00502	LYIOS00502	LYIOS00502	LYIOS00503	LYIOS00503	LYIOS00503	LYIOS00503

The following science campaigns were performed by LYRA:

- Daily U3 observation campaigns
- Short bi-weekly calibration on 2015-Oct-21
- Between 2015-Oct-23 20:48 and 2015-Oct-24 02:22 the visual LED was switched on (uncommanded)

### LYRA detector temperature

LYRA detector 2 temperature globally varied between 49.66 and 52.52 °C.

### 3. SWAP instrument status

### Calibration

Calibration campaign on Tuesday this week.

#### **MCPM errors**

The number of MCPM recoverable errors increased from 174 to 316. The number of MCPM unrecoverable errors remained 0.

### **IOS & operations**

Monday	Tuesday	Wednesday	Thursday	Friday	Saturday	Sunday
19 Oct	20 Oct	21 Oct	22 Oct	23 Oct	24 Oct	25 Oct
Nominal acquisition	Nominal acquisition + calibration	Nominal acquisition				
IOS00602	IOS00602	IOS00602	IOS00602	IOS00602	IOS00602	IOS00602
597 images	693 images	570 images	680 images	640 images	521 images	570 images

Special operations for SWAP, this week:

• Bi-weekly calibration on 2015-Oct-20

### **SWAP** detector temperature

The SWAP Cold Finger Temperature globally varied between 1.83 and 2.87 °C.

## 4. PROBA2 Science Center Status

The main operator is Katrien Bonte.

The following changes were made to the P2SC:

• None.

## 5. Data reception & discussions with MOC

#### **Passes**

The delivery of the passes for this week (passes 18763 to 18824) was nominal, except for:

None.

### Data coverage HK

All HK data files (LYRA AD) have been received, except:

None.

### **Data coverage SWAP**

All SWAP Science data files (BINSWAP) have been received, except:

None.

Total number of images between 2015 Oct 19 0UT and 2015 Oct 26 0UT: 4272

Highest cadence in this period: 0 seconds

Average cadence in this period: 141.59 seconds Number of image gaps larger than 300 seconds: 205

Largest data gap: 11.00 minutes

### **Data coverage LYRA**

All LYRA Science data files (BINLYRA) have been received, except:

None

## 6. APPENDIX: Frequently used acronyms

ADPMS Advanced Data and Power Management System

AOCS Attitude and Orbit Control System

APS Active Pixel image Sensor

ASIC Application Specific Integrated Circuit

BBE Base Band Equipment CME Coronal Mass Ejection

COGEX Cool Gas Generator Experiment

CRC Cyclic Redundancy Check
ESP Experimental Solar Panel

FITS Flexible Image Transport System

FOV Field Of View FPA Focal Plane Assembly

FPGA Field Programmable Gate Arrays

GPS Global Positioning System

HK Housekeeping

IOS Instrument Operations Sheet

LED Light Emitting Diode
LYRA LYman alpha RAdiometer

LYTMR LYRA Telemetry Reformatter (software module of P2SC)
LYEDG LYRA Engineering Data Generator (software module of P2SC)

MCPM Mass Memory, Compression and Packetisation Module

MOC Mission Operation Center NDR Non Destructive Readout

OBSW On board Software
PI Principal Investigator
P2SC PROBA2 Science Center
ROB Royal Observatory of Belgium

SAA South Atlantic Anomaly SEU Single Event Upset

SoFAST | Solar Feature Automated Search Tool

SWAP Sun Watcher using APS detector and image Processing

SWAVINT | SWAP AVerage INTensity

SWBSDG SWAP Base Science Data Generator

SWEDG SWAP Engineering Data Generator (software module of P2SC)
SWTMR SWAP Telemetry Reformatter (software module of P2SC)

TBC To Be Confirmed
TBD To Be Defined
TC Telecommand

UTC Coordinated Universal Time

UV Ultraviolet

VFC Voltage to Frequency Converter

# 7. APPENDIX Solar Activity Definitions

In the science section we use the following solar activity standards.

The standard scale for solar activity is:

- very low (almost no flares, only B)
- low (a few C flares)
- moderate (many C flares and at least an M flare)
- high (several M flares and an X flare)
- very high (continuous background of C flares, numerous M flares, more than one X flare)