P2SC-ROB-WR-287 - 20150921 Weekly report #287	P2SC Weekly report	**** <u>***</u>
Period covered: Date:	Mon Sep 21 to Sun Sep 27, 2015 30 Sep 2015	Royal Observatory of Belgium -
Written by: Approved by:	Katrien Bonte Matthew West	PROBA2 Science Center
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# 1. Science

# Solar & Space weather events

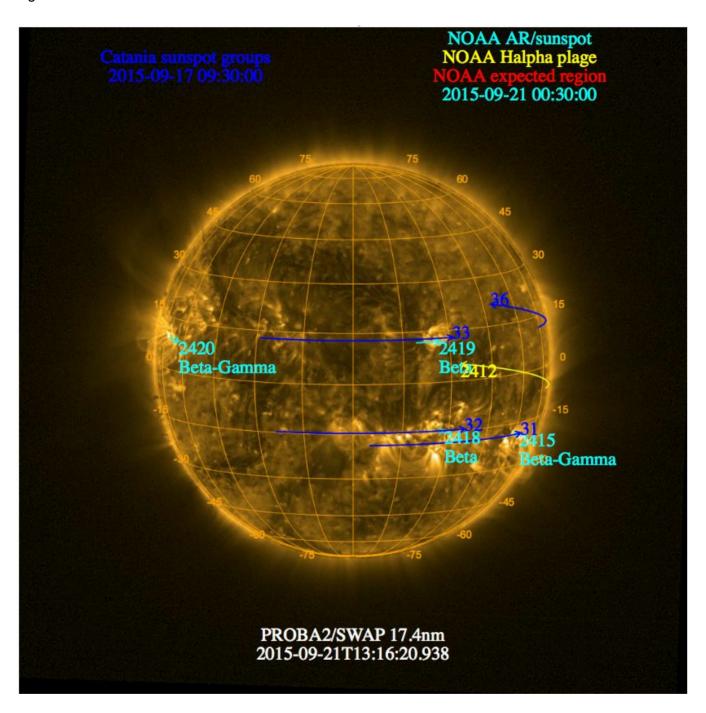
The level of solar activity¹ fluctuated between **very low** and **moderate** this week.

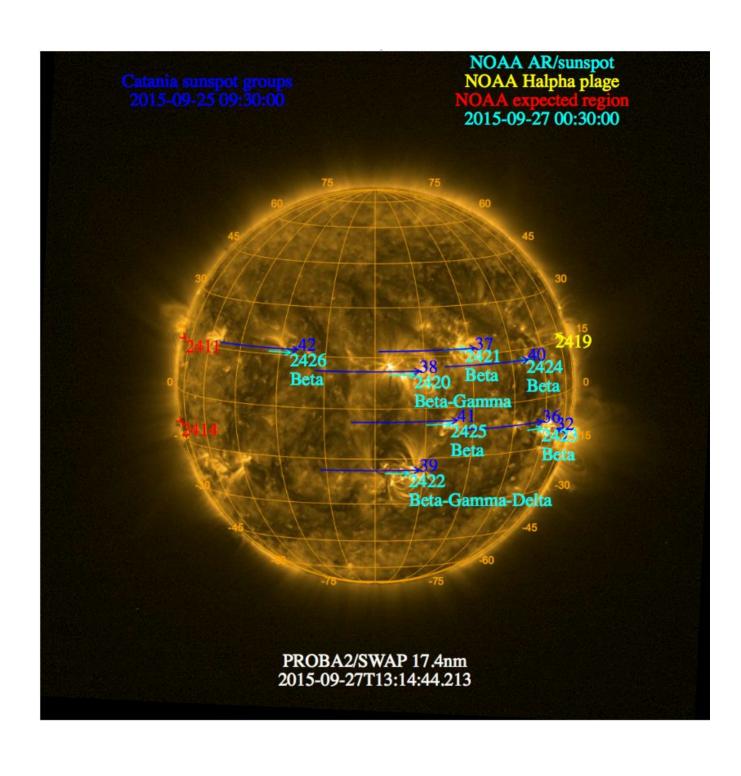
Only M- and X-flares are mentioned, the most energetic one(s) per day are presented in **bold**:

	Monday 21 Sep	Tuesday 22 Sep	Wednesday 23 Sep	Thursday 24 Sep	Friday 25 Sep	Saturday 26 Sep	Sunday 27 Sep
Activity	low	very low	low	low	low	low	moderate
Flares	-	-	-	-	-	-	M1.9@10:40 M1.0@21:00

<sup>&</sup>lt;sup>1</sup> See appendix. All timings are given in UT.

The SWAP images of 2015-09-21 and Sep 2015-09-27 are shown below, with annotated active regions.





## **Solar Activity**

Solar flare activity fluctuated between very low and moderate during the week. In order to view the activity of this week in more detail, we suggest to go to the following website from which all the daily (normal and difference) movies can be accessed: <a href="http://proba2.oma.be/ssa">http://proba2.oma.be/ssa</a>
This page also lists the recorded flaring events.

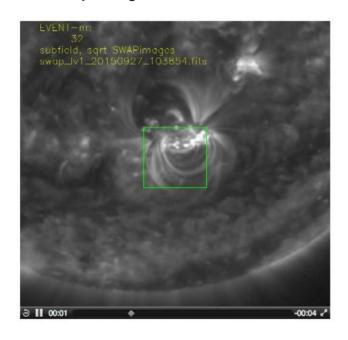
A weekly overview movie can be found here (SWAP week 287).

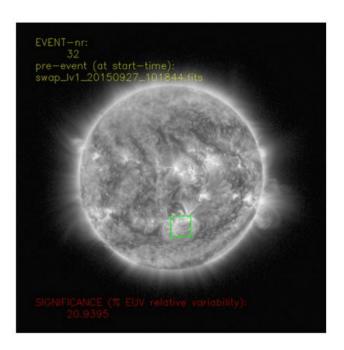
Details about some of this week's events:

Solar activity increased towards the end of the week. AR 2422 produced two M-class flares on 2015-09-27: an M1.9 flare peaking around 10:40 UT and an M1.0 flare peaking around 21:00 UT. Below we provide SWAP images from the time when these M-flares occurred on 2015-Sep-27. The annotated snapshots are produced by the Solar Feature Automated Search Tool (SoFAST). This tool detects dynamic solar events in EUV images from SWAP in near real-time. The snapshots illustrate the location of the flare on the solar disk (right) and a zoomed image (left).

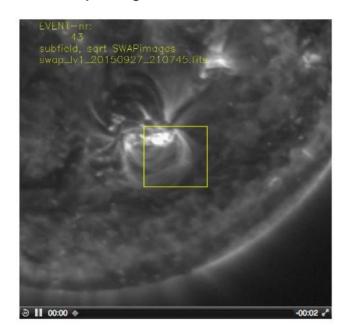
The complete SoFAST online event list and additional plots are available on <a href="http://www.sidc.be/sofast">http://www.sidc.be/sofast</a>.

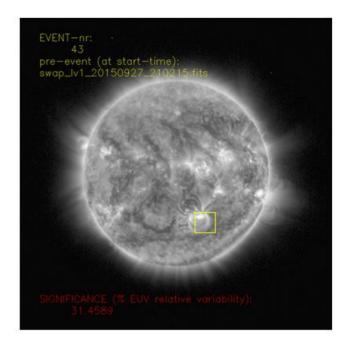
# 2015-09-27: M1.9 flare peaking around 10:40 UT





# M1.0 flare peaking around 21:00 UT

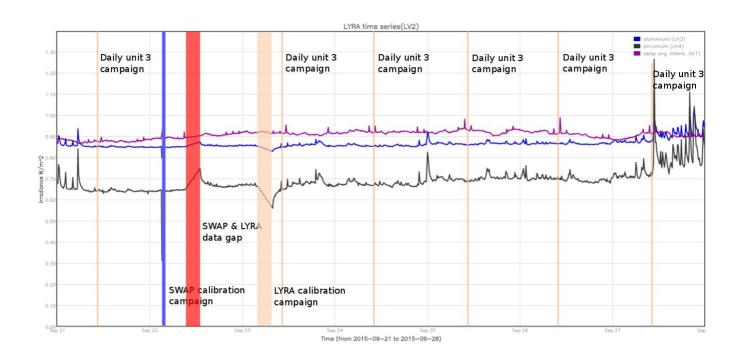




An overview of the weekly LYRA & SWAP data is provided below:

#### The following curves are visible:

- black: Zirconium Channel LYRA Unit 2
- blue: Aluminium Channel of LYRA Unit 2
- purple: SWAVINT (SWAP Average Intensity; integrated solar intensity per SWAP image pixel)



# The blue shaded periods correspond to, from left to right:

• SWAP bi-weekly calibration campaign on 2015-09-22

## The orange shaded periods correspond to, from left to right:

- LYRA daily U3 campaign on 2015-09-21
- LYRA short bi-weekly calibration on 2015-09-23
- LYRA daily U3 campaign on 2015-09-23
- LYRA daily U3 campaign on 2015-09-24
- LYRA daily U3 campaign on 2015-09-25
- LYRA daily U3 campaign on 2015-09-26
- LYRA daily U3 campaign on 2015-09-27

## The red shaded periods correspond to:

SWAP and LYRA data gap due to spacecraft in safe mode

#### Outreach, papers, presentations, etc.

Please consult <a href="http://proba2.oma.be/science/publications">http://proba2.oma.be/science/publications</a> for a list of interesting articles using SWAP & LYRA data, as well as a link to the complete article list.

The science section of this weekly report is also published in the weekly STCE newsletter (http://www.stce.be/newsletter/newsletter.php).

#### Meeting

The CNES, the LATMOS and the STCE organised the second edition of the meeting **"Solar Metrology, Needs and Methods"**. This meeting was open to anyone involved in solar metrology and modelling.

It took place from 2015-09-21 to 2015-09-23, at the Royal Observatory of Belgium.

PROBA2 related talks that were on the program include:

Mid-term Periodicities of the LYRA data spectrum

L. Wauters

Long-term variability of LYRA data

I. Dammasch

Space-based instrument developments for UV solar observations - detector technology - A. Benmoussa

Progress towards understanding the degradation affecting the PROBA2/LYRA instrument M. Dominique

## **Guest Investigator Program**

None

# 2. LYRA instrument status

#### Calibration

Calibration campaign on Wednesday this week.

# **IOS & operations**

Monday 21 Sep	Tuesday 22 Sep	Wednesday 23 Sep	Thursday 24 Sep	Friday 25 Sep	Saturday 26 Sep	Sunday 27 Sep
Nominal acquisition + daily U3	Nominal acquisition	Nominal acquisition + daily U3 + calibration	Nominal acquisition + daily U3	Nominal acquisition + daily U3	Nominal acquisition + daily U3	Nominal acquisition + daily U3
LYIOS00494	LYIOS00494 -> LYIOS00495	LYIOS00495	LYIOS00495	LYIOS00496	LYIOS00496 -> LYIOS00497	LYIOS00497

The following science campaigns were performed by LYRA:

- Daily U3 observation campaigns, except on 2015-09-22
- Short bi-weekly calibration on 2015-09-23

# LYRA detector temperature

LYRA detector 2 temperature globally varied between 43.59 and 49.64 °C.

# 3. SWAP instrument status

#### Calibration

Calibration campaign on 2015-09-22.

## **MCPM** errors

The number of MCPM recoverable errors remained 137.

The number of MCPM unrecoverable errors remained 0.

# **IOS & operations**

Monday 21 Sep	Tuesday 22 Sep	Wednesday 23 Sep	Thursday 24 Sep	Friday 25 Sep	Saturday 26 Sep	Sunday 27 Sep
Nominal acquisition	Nominal acquisition + calibration	Nominal acquisition				
IOS00597	IOS00597 ->IOS00598	IOS00598	IOS00598	IOS00599	IOS00599	IOS00599
653 images	569 images	528 images	672 images	574 images	552 images	483 images

Special operations for SWAP, this week:

• Bi-weekly calibration on 2015-09-22

# **SWAP** detector temperature

The SWAP Cold Finger Temperature globally varied between -4.25 and -0.39 °C.

# 4. PROBA2 Science Center Status

The main operator is Katrien Bonte.

The following changes were made to the P2SC:

• None.

# 5. Data reception & discussions with MOC

#### **Passes**

The delivery of the passes for this week (passes 18526 to 18582) was nominal, except for:

None.

## Data coverage HK

All HK data files (LYRA AD) have been received, except:

None.

# **Data coverage SWAP**

All SWAP Science data files (BINSWAP) have been received, except:

None.

Total number of images between 2015-09-21 00:00 UT and 2015-09-28 00:00 UT: 3378

Highest cadence in this period: 30 seconds Average cadence in this period: 153.48 seconds Number of image gaps larger than 300 seconds: 197

Largest data gap: 222.95 minutes

There is a significant data gap because the spacecraft went to safe mode on 2015-09-22.

# **Data coverage LYRA**

All LYRA Science data files (BINLYRA) have been received, except:

None.

# 6. APPENDIX: Frequently used acronyms

ADPMS Advanced Data and Power Management System

AOCS Attitude and Orbit Control System

APS Active Pixel image Sensor

ASIC Application Specific Integrated Circuit

BBE Base Band Equipment
CME Coronal Mass Ejection

COGEX Cool Gas Generator Experiment

CRC Cyclic Redundancy Check ESP Experimental Solar Panel

FITS Flexible Image Transport System

FOV Field Of View FPA Focal Plane Assembly

FPGA Field Programmable Gate Arrays

GPS Global Positioning System

HK Housekeeping

IOS Instrument Operations Sheet

LED Light Emitting Diode
LYRA LYman alpha RAdiometer

LYTMR LYRA Telemetry Reformatter (software module of P2SC)
LYEDG LYRA Engineering Data Generator (software module of P2SC)

MCPM Mass Memory, Compression and Packetisation Module

MOC Mission Operation Center NDR Non Destructive Readout

OBSW On board Software
PI Principal Investigator
P2SC PROBA2 Science Center
ROB Royal Observatory of Belgium

SAA South Atlantic Anomaly SEU Single Event Upset

SoFAST | Solar Feature Automated Search Tool

SWAP Sun Watcher using APS detector and image Processing

SWAVINT | SWAP AVerage INTensity

SWBSDG SWAP Base Science Data Generator

SWEDG SWAP Engineering Data Generator (software module of P2SC)
SWTMR SWAP Telemetry Reformatter (software module of P2SC)

TBC To Be Confirmed
TBD To Be Defined
TC Telecommand

UTC Coordinated Universal Time

UV Ultraviolet

VFC Voltage to Frequency Converter

# 7. APPENDIX Solar Activity Definitions

In the science section we use the following solar activity standards.

The standard scale for solar activity is:

- very low (almost no flares, only B)
- low (a few C flares)
- moderate (many C flares and at least an M flare)
- high (several M flares and an X flare)
- very high (continuous background of C flares, numerous M flares, more than one X flare)